Development of line removal methods

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Outline

Introduction

- What's line (narrow band noise)
- Motivation

Line removal method->MBLT

- Method
- implementation

Operation check

- various Q values
- effect to the detection of GW burst signal
- apply to real data(LIGO S5)

Summary and future plan

Introduction

What's narrow band noise

KAGRA's design sensitivity curve



Narrow band noise of GW detector (called line)

-> Caused by the elastic vibration of the wire of the noise and the suspension system of power

Motivation

Problem of Lines

The detection of GW signal becomes difficult when the overlap in terms of frequency narrow band noise
Lines in high sensitivity area



Important to remove the lines with leaving GW signal
Clarify what kind of lines can be removed
Clarify the benefits of the detection by line removal

Line removal methods

LIGO's past observations

When the line affecting the signals, signals was inhibited with lines

Line Removal method that was used this time

Soumya D Mohanty (2002)stacks.iop.org/ CQG/19/1513

MBLT(Median Based Line Tracker)

< Feature >

- I small effect on the transient signal
- possible to follow the frequency variation of lines
- remove lines in time domain
- □ depends on the center frequency and Q value of lines

MBLT Algorithm

Assume that center frequency and Q values of lines are known



Repeat this steps until removing lines

<u>Center frequency and Q value are important</u> (ex : change the accuracy of the median estimated by block size)

MBLT operation check



MBLT operation check

1.5

0.5

0

-0.5

-1.5

-2

-2.5 2.5

1.5

0.5

0

-0.5

-1

-2

-2.5

-1.5



Motivation

Important to remove the lines with leaving GW signal
Clarify what kind of lines can be removed

Clarify the benefits of the detection by line removal

Is it possible to remove the line while leaving GW burst signal ?

Effect on GW burst analysis

KAGRABurstPlusBefore

Injection GW signal and line to KAGRA's simulation data



For real data

◆LIGO's data (LIGO S5(H1) data) T=10[s],fs=2048[Hz]

LIGO data which was conditioned other than line



the lines in 17,35,60,337,345,695Hz

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For real data



summary and future plan

summary

- •We have implemented MBLT
- •High Q lines are advantageous when removing with MBLT
- Success to remove line with leaving signal
- •Operation check by real data(LIGO S5)
- -> 100Hz~ lines can be removed successfully
- -> ~100Hz lines cannot be removed -> need more detail checks

future plan

•Measures in the case where there is more than one line at a frequency close

- Understand what type of lines can remove or can't remove
 - -> Q value dependence
 - -> frequency dependence
 - -> noise level dependence

Reduce the lines that cannot be removed

