

Developing a novel method for finding correlated features between control, environment channels

DetChar Meeting
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Hiroataka Yuzurihara
Osaka City University

Collaborator : Kazuhiro Hayama, Shuhei Mano,
Takahiro Yamamoto, Atsushi Nishizawa, Nobuyuki Kanda

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- Remove false trigger event in GW search

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- Maximum Information Coefficient

Developed Monitor Tool

- Developed correlation map
- Analysis result to search correlation between accelerometer and GW channel

motivation - correlation analysis using environmental channels

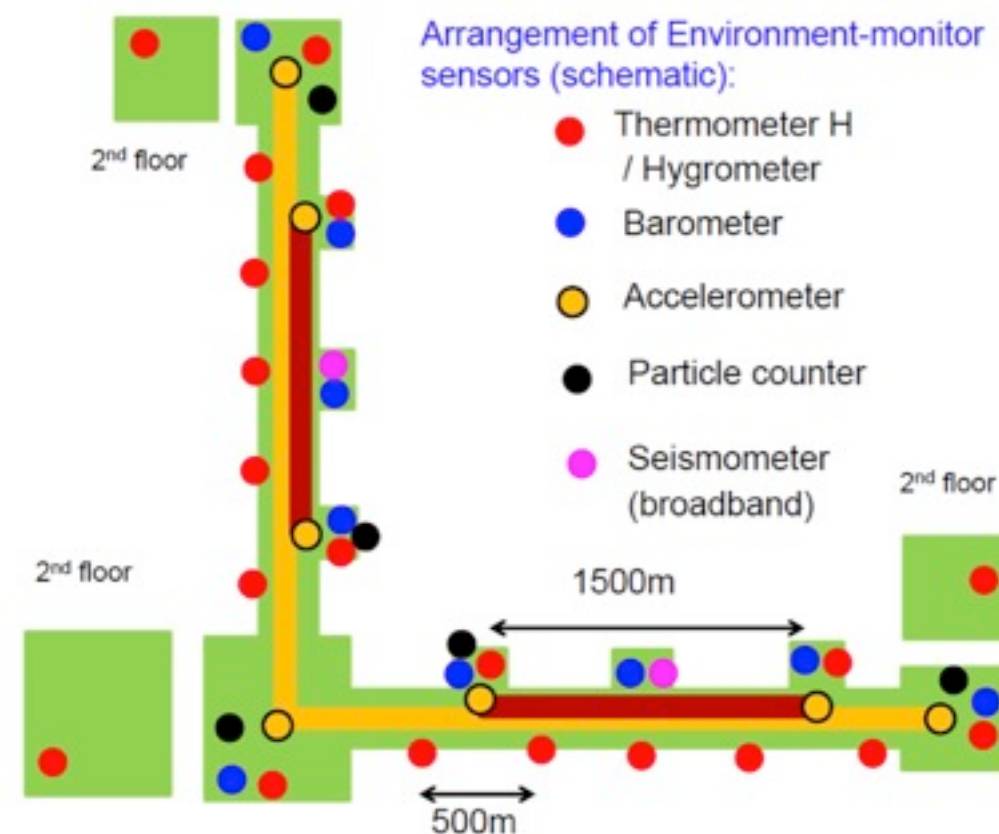
Goal : search correlated channels between ~10000 environmental channels and finally localize noise sources

- localization of noise sources reveals noise features
- remove false trigger event generated by GW search pipeline,
 - > increase GW detection efficiency

In this talk, I define

- GW channel
as sensitive channel to GW
- environmental channel
as insensitive channel to GW

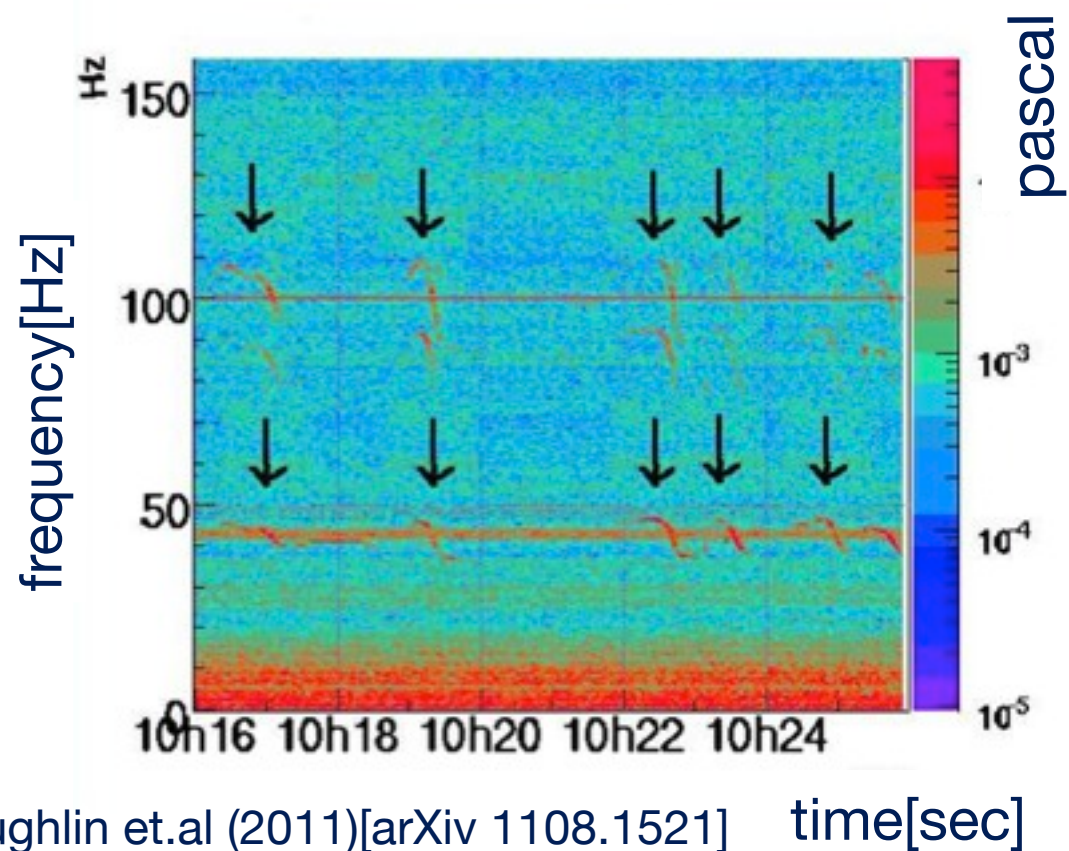
microphone, accelerometer,
seismometer, thermometer,
barometer, magnetometer
so on



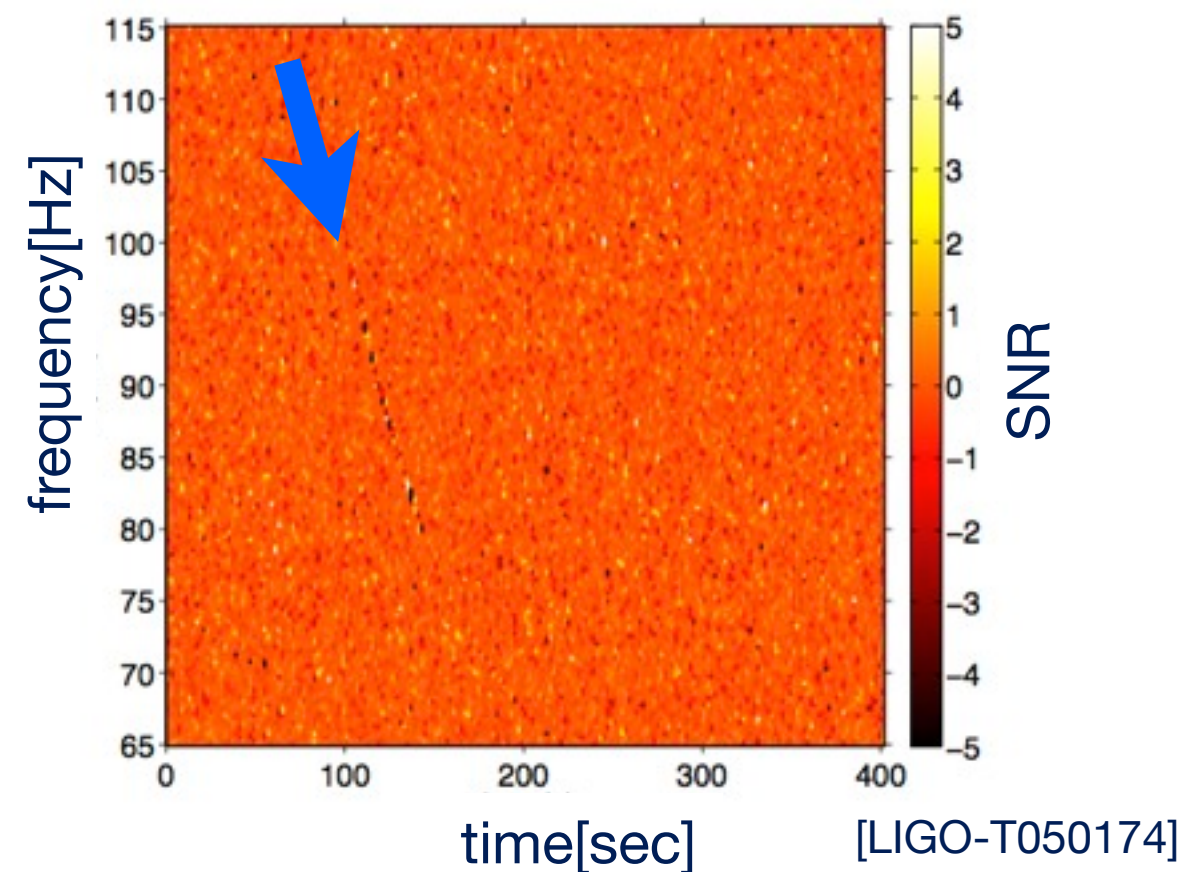
Araya (2012)

Example of linear correlation observed in LIGO and Virgo

Helicopter

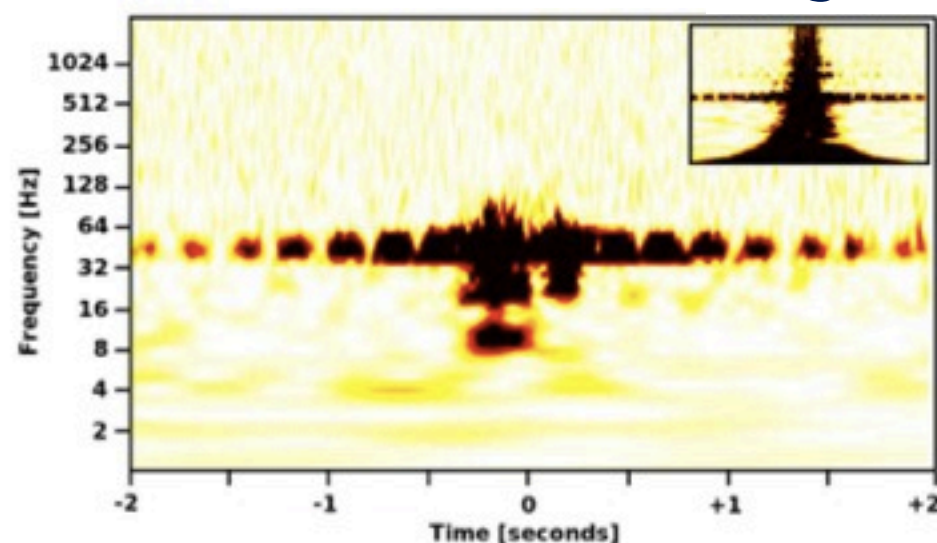


Airplane



GW channel

Magnetic sensor



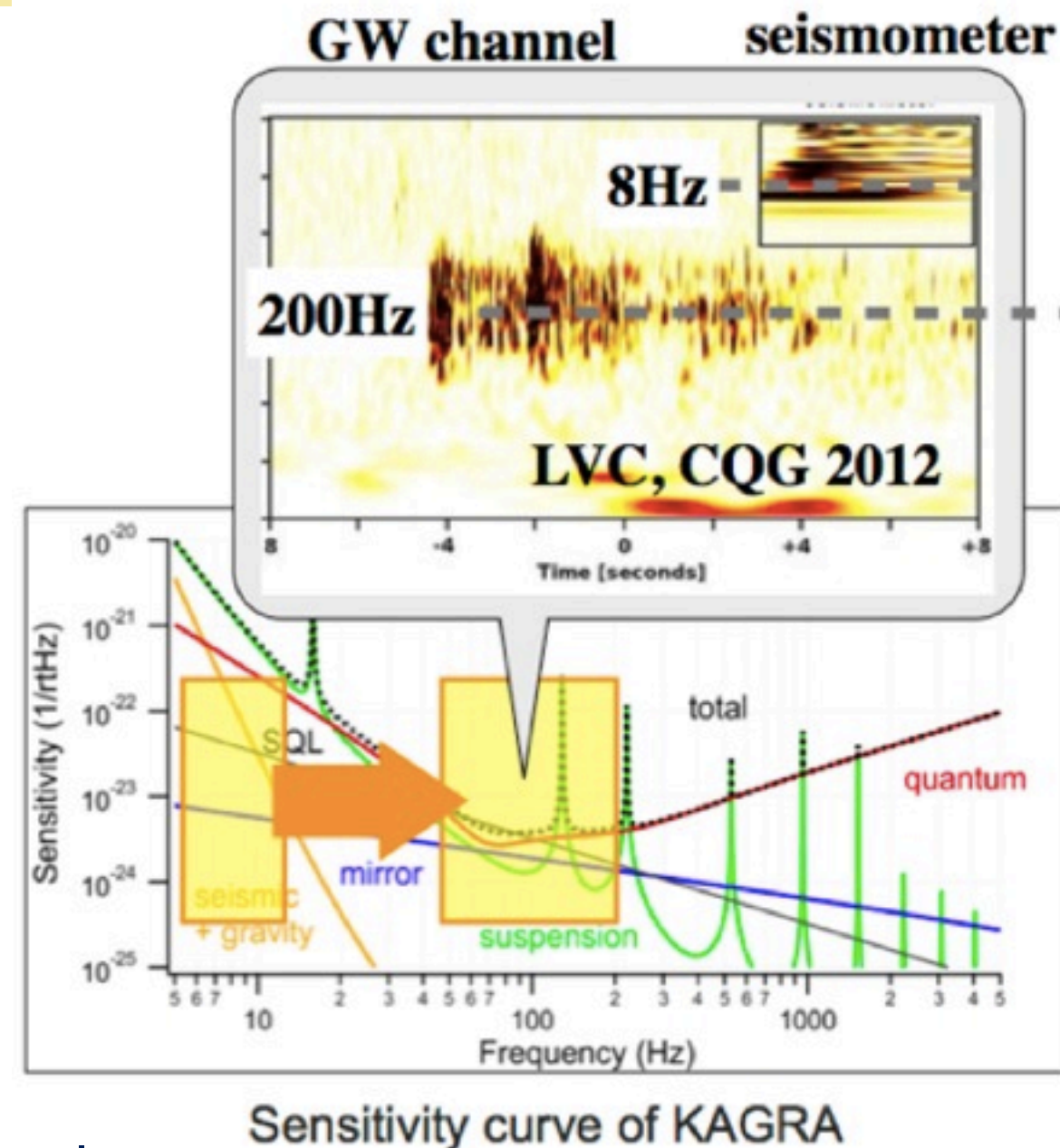
J.Asis et al. (2012) [gr-qc 1203.5613]

Example of non-linear correlation, up-converted noise

Up-converted noise: seismic glitches will excite optical bench motion which cause scattered light noise. In bad weather day, effect from seismic glitches is strong. Non-linear correlation over a few Hz ~ a few hundreds Hz in GW channel was observed in detectors.

[Classical and Quantum Gravity 27, 19 (2010) 194011]

<http://www.opticsinfobase.org/oe/abstract.cfm?uri=oe-20-8-8329>



⇒ search correlation between environmental channels
to get information of noise which effect different frequency band

Hayama (2014)

J.Asis et al. (2012) [gr-qc 1203.5613]

Developing monitor tool to search linear & non-linear correlation

- In KAGRA detector commissioning and observation period, efficient tools to localize noises
- Developed software is included in HasKAL
HasKAL is Haskell-based analysis software package.
You can easily access from below URL.

<https://github.com/gw-analysis/detector-characterization>

open access @GitHub



working with lpef		latest commit f1c75831f2
enoshima authored 10 hours ago		
HasKAL	working with lpef	10 hours ago
attic	working with lpef	10 hours ago
optFiles	minor fix	2 days ago
test	minor fix	2 days ago
.gitignore	modified HasKAL.cabal	9 days ago

Method to search correlation

- In this study, below two method are used,

Pearson Correlation Coefficient

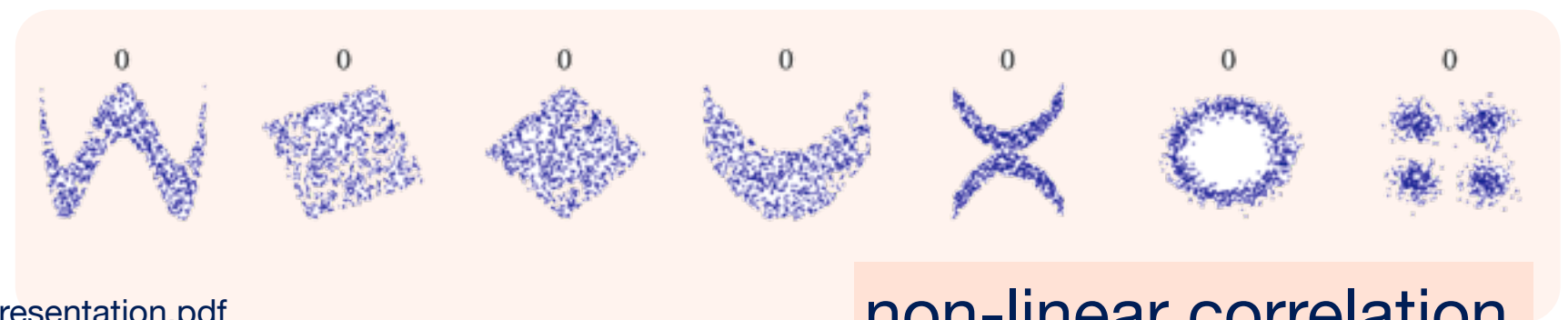
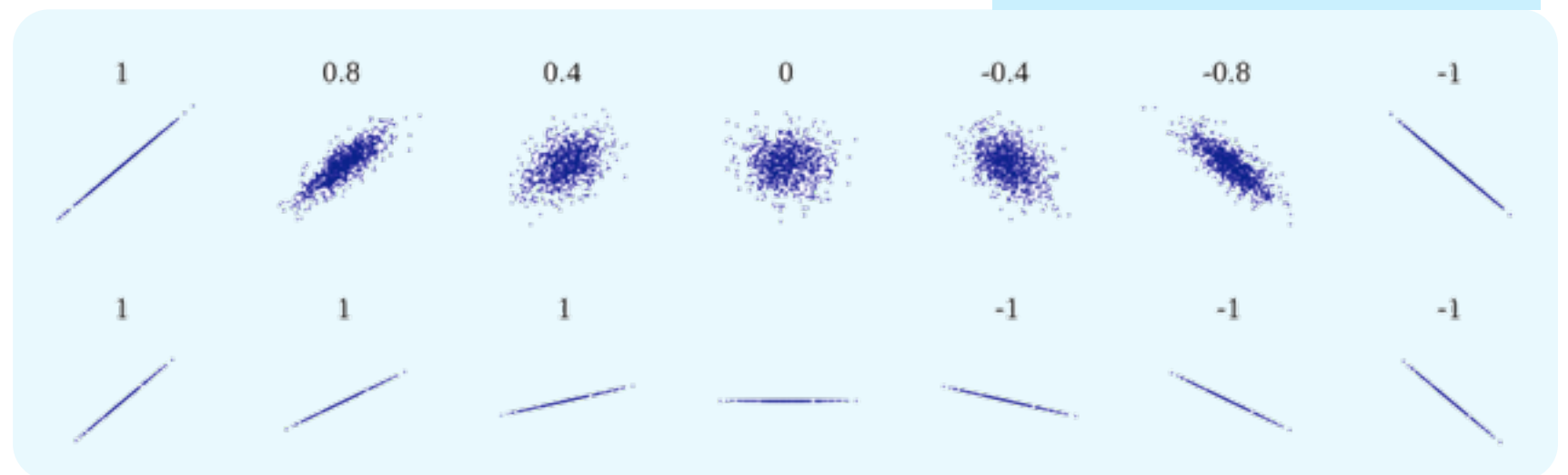
- efficient method to linear correlation

$$r = \frac{\sum_i (x_i - \bar{x})(y_i - \bar{y})}{\sqrt{\sum_i (x_i - \bar{x})^2 \sum_i (y_i - \bar{y})^2}}$$

Maximum Information Coefficient (MIC)

- structure of scatter plot
→ Information of
noise transfer

linear correlation



non-linear correlation

Maximum Information Coefficient (MIC)

[David N. Reshef, et al. Science 334, 1518 (2011)]

- MIC can detect functional & non-functional dependence.
- If provided sufficient sample size, MIC can detect a wide range of relationship including non-functional types (functional superposition)
- If a relationship exists between two data, a grid can be drawn on the scatter plot of two data that partitions the data to encapsulate that relationship.

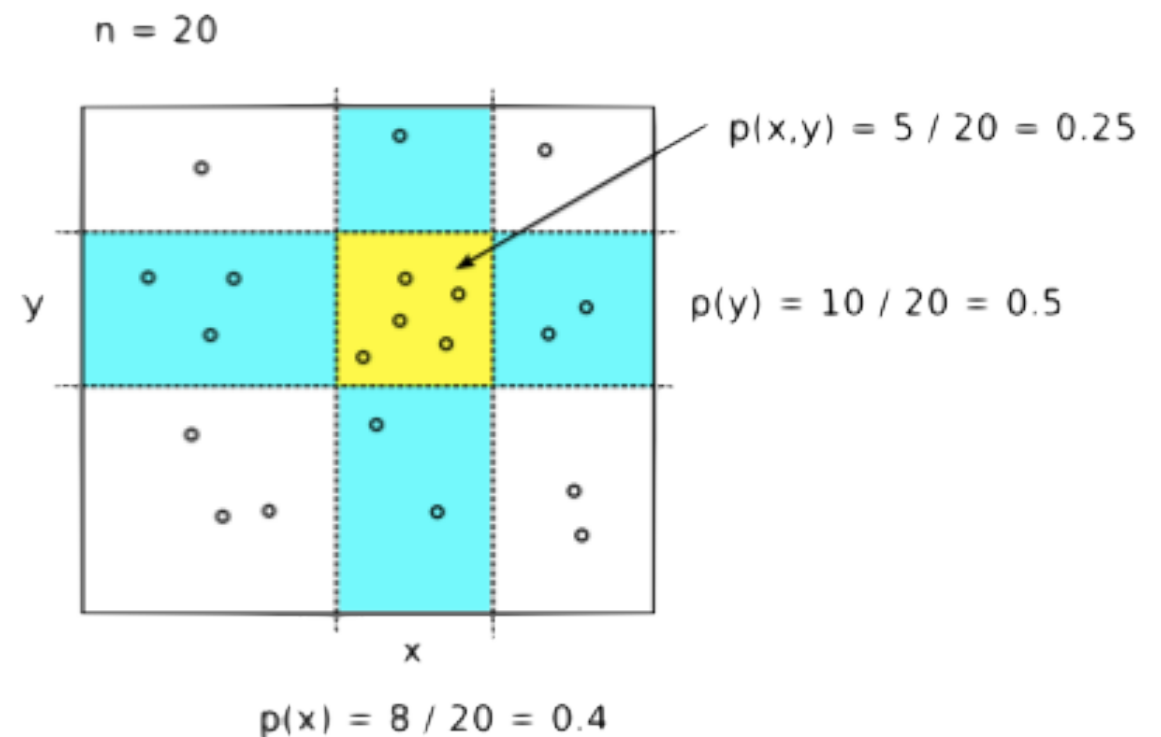
Partitioning : For each resolution, MIC finds grid partition placement with highest mutual information

$$I(X; Y) = \sum_{y \in Y} \sum_{x \in X} p(x, y) \log \left(\frac{p(x, y)}{p(x) p(y)} \right)$$

X, Y: random variables

$p(x, y)$: joint probability distribution function

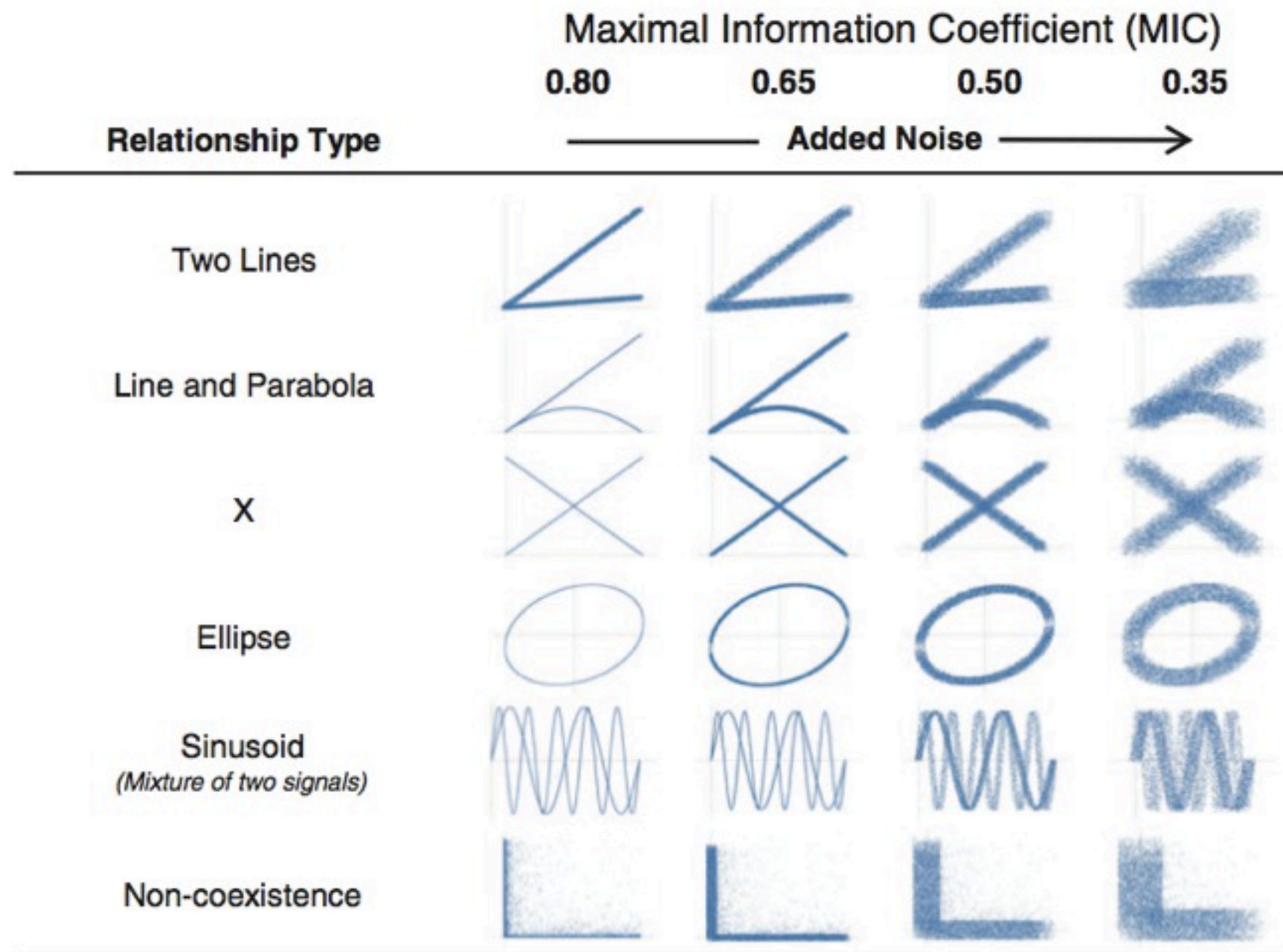
$p(x)$, $p(y)$: marginal probability distribution functions



$$p(x, y) \log \left(\frac{p(x, y)}{p(x) p(y)} \right) = 0.25 \log \left(\frac{0.25}{0.4 \times 0.5} \right) \approx 0.056$$

Which correlation MIC can search?

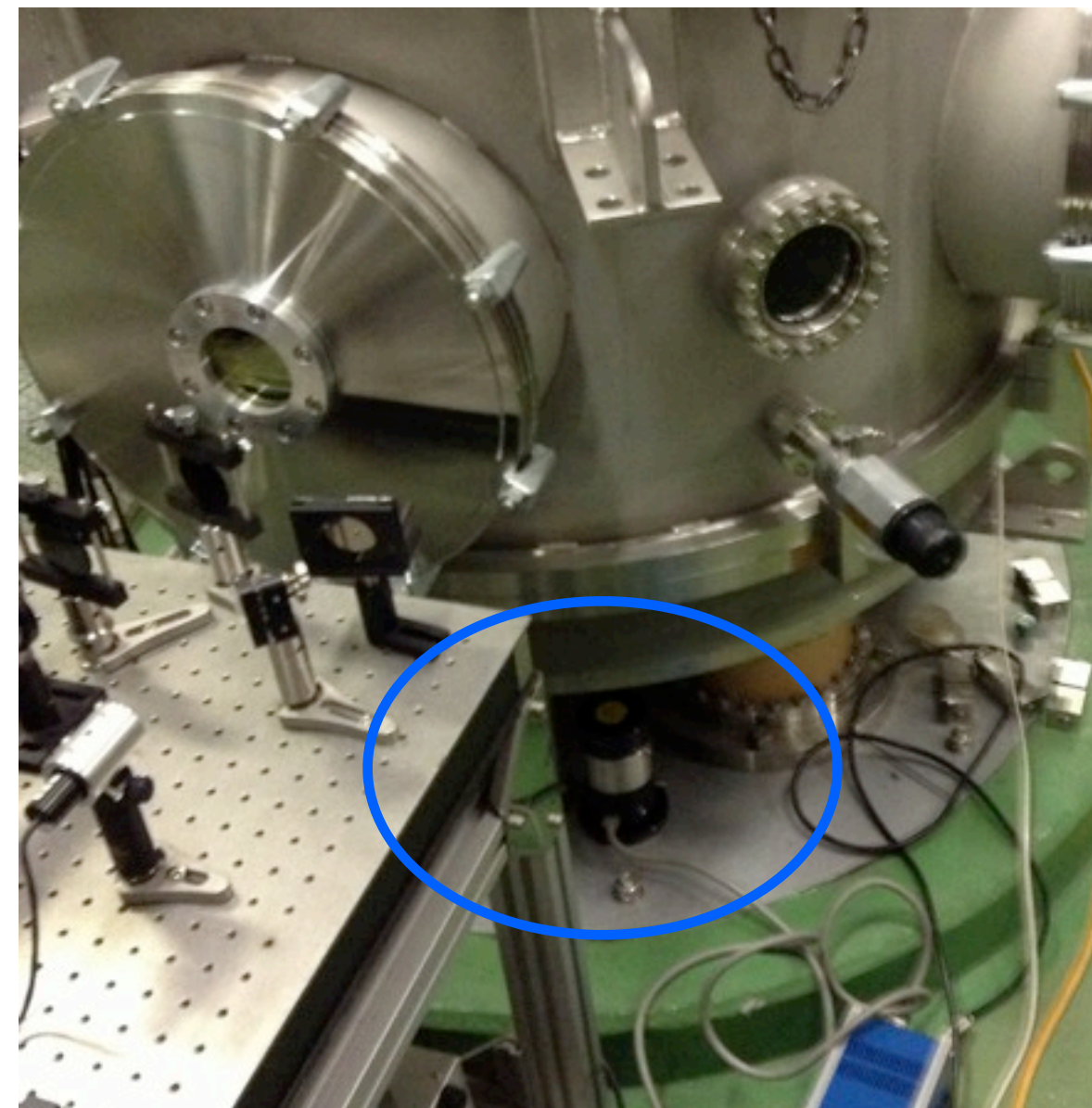
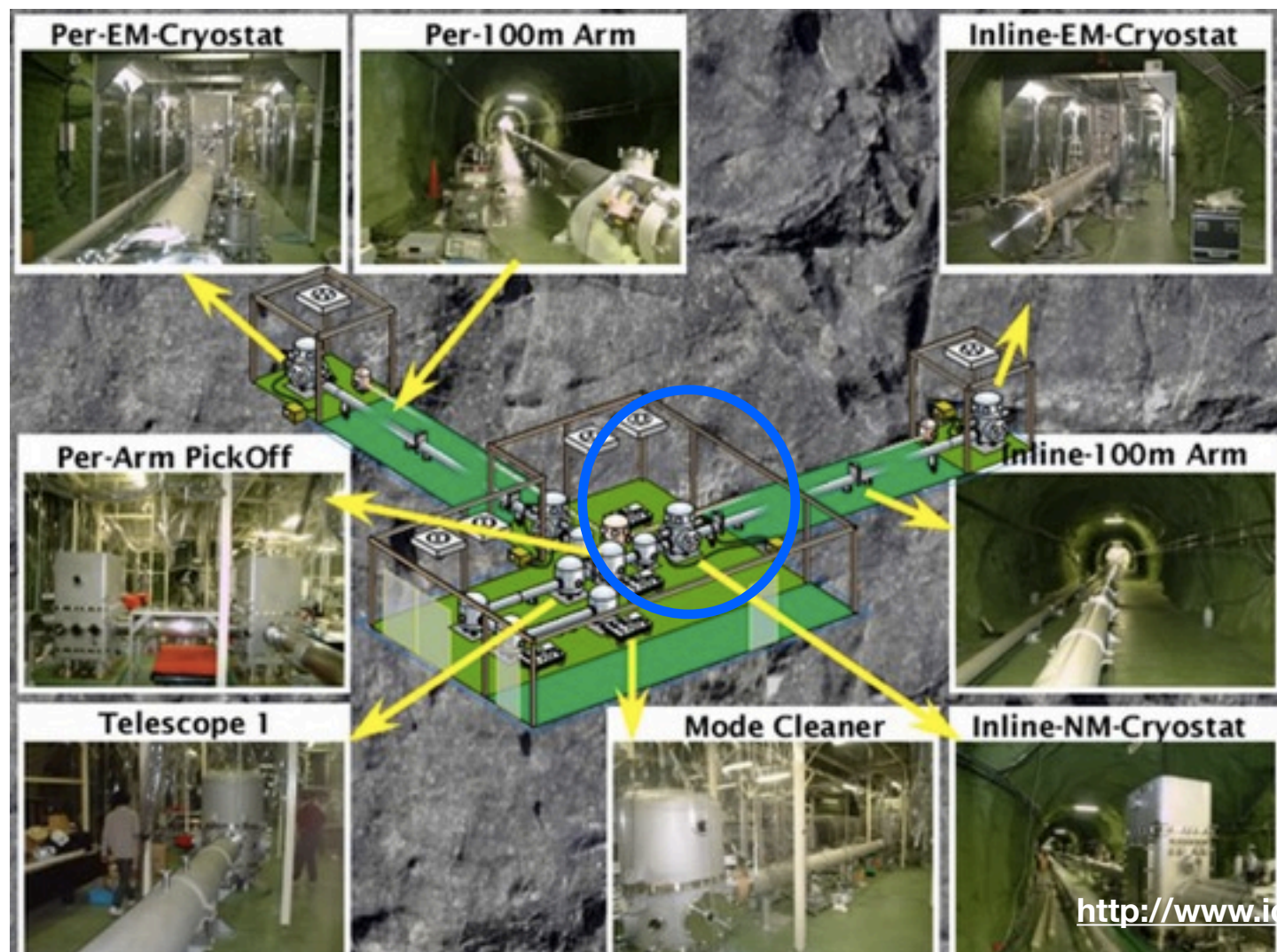
- MIC can search not only linear but also non-linear correlation.



Operation check using CLIO's channel data

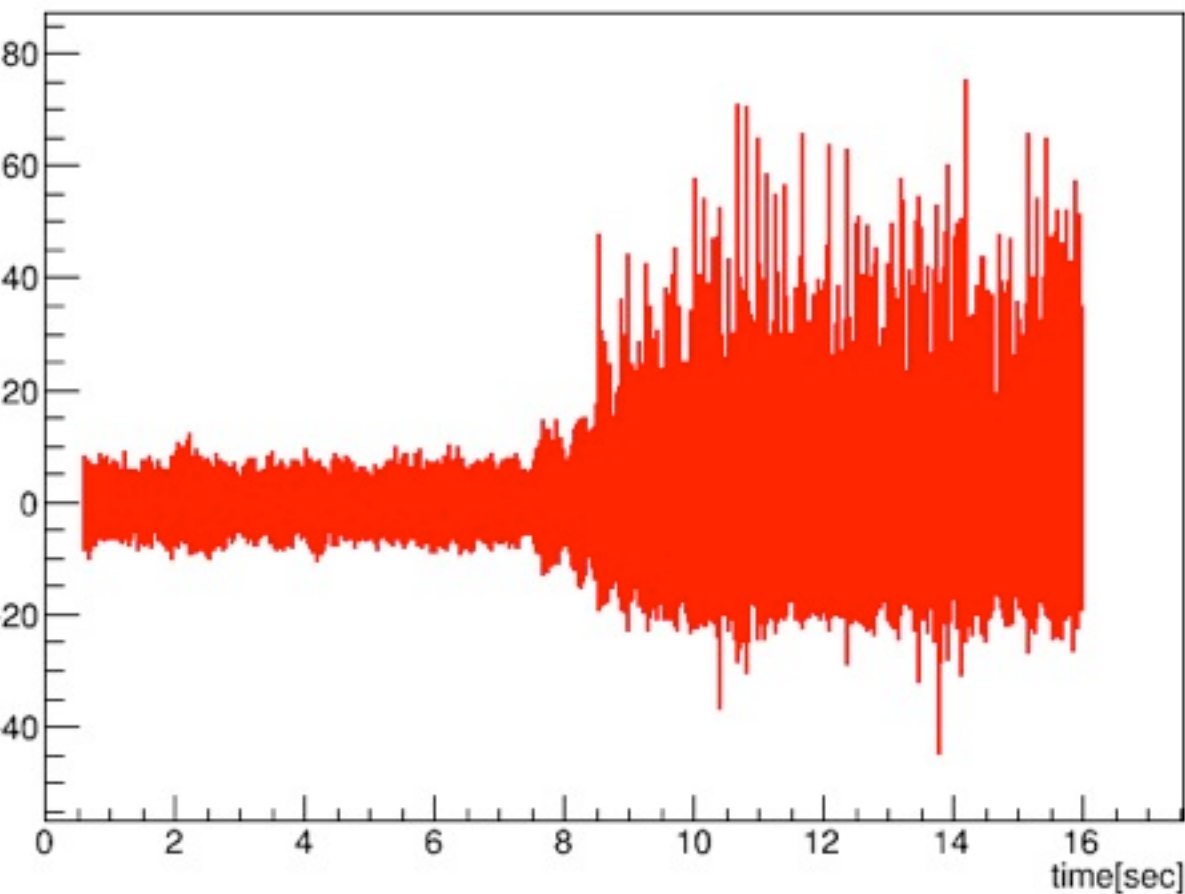
- CLIO detector data recorded at September 2012
two channel data : GW channel
accelerometer channel
- search for correlation between GW channel and accelerometer channel

Accelerometer is located near front mirror of X-arm.



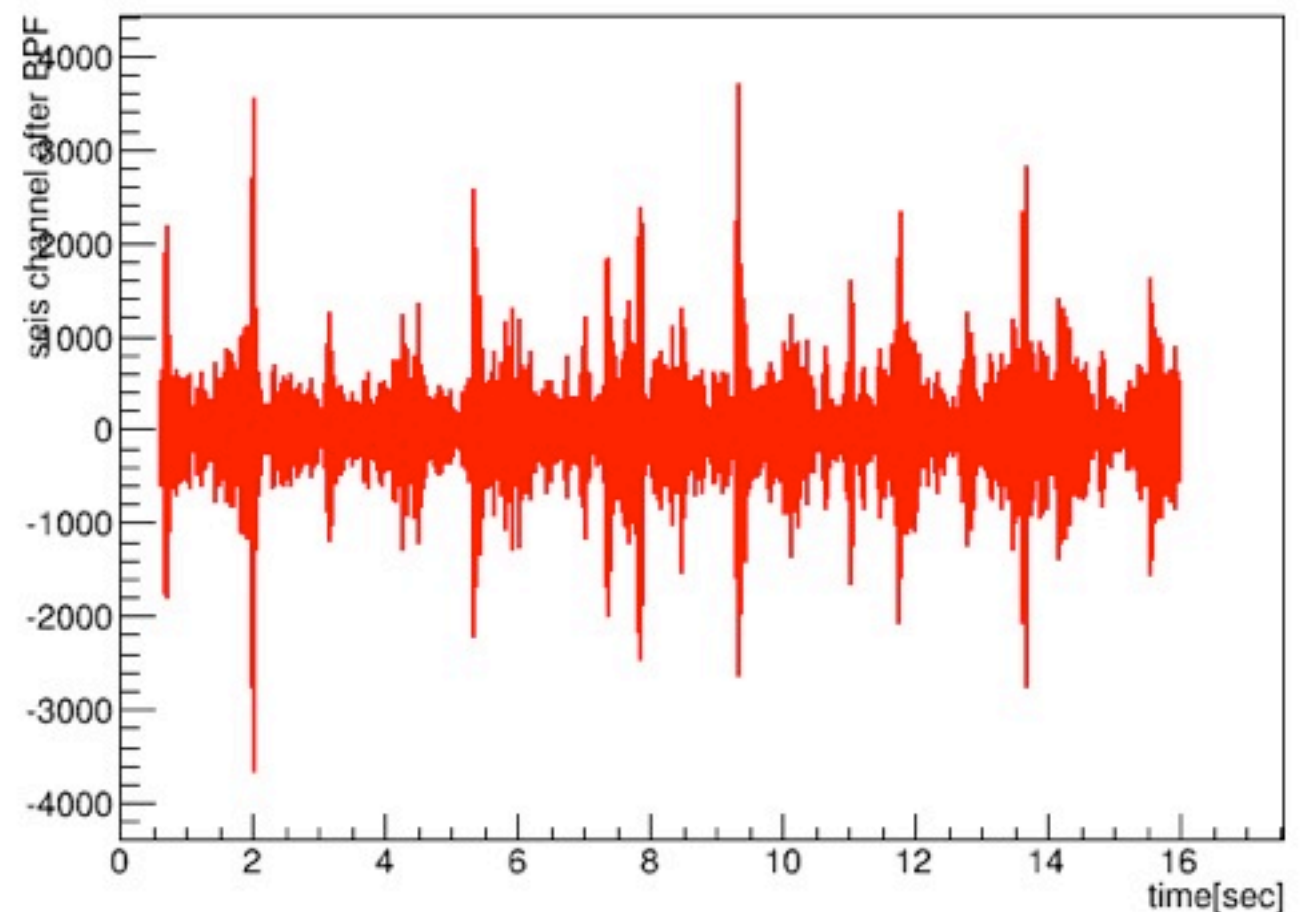
time domain data used to search

1034657664



GW channel

1034657664

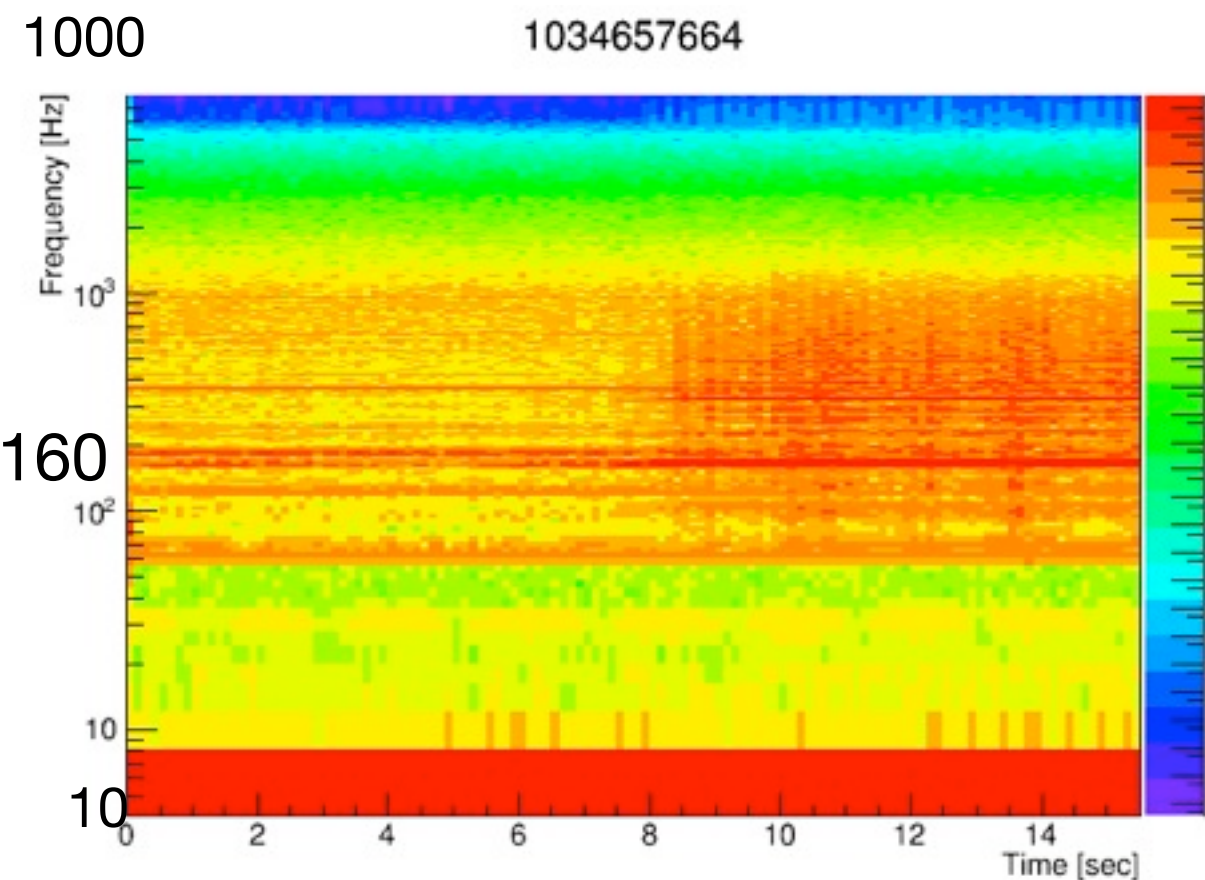


accelerometer channel

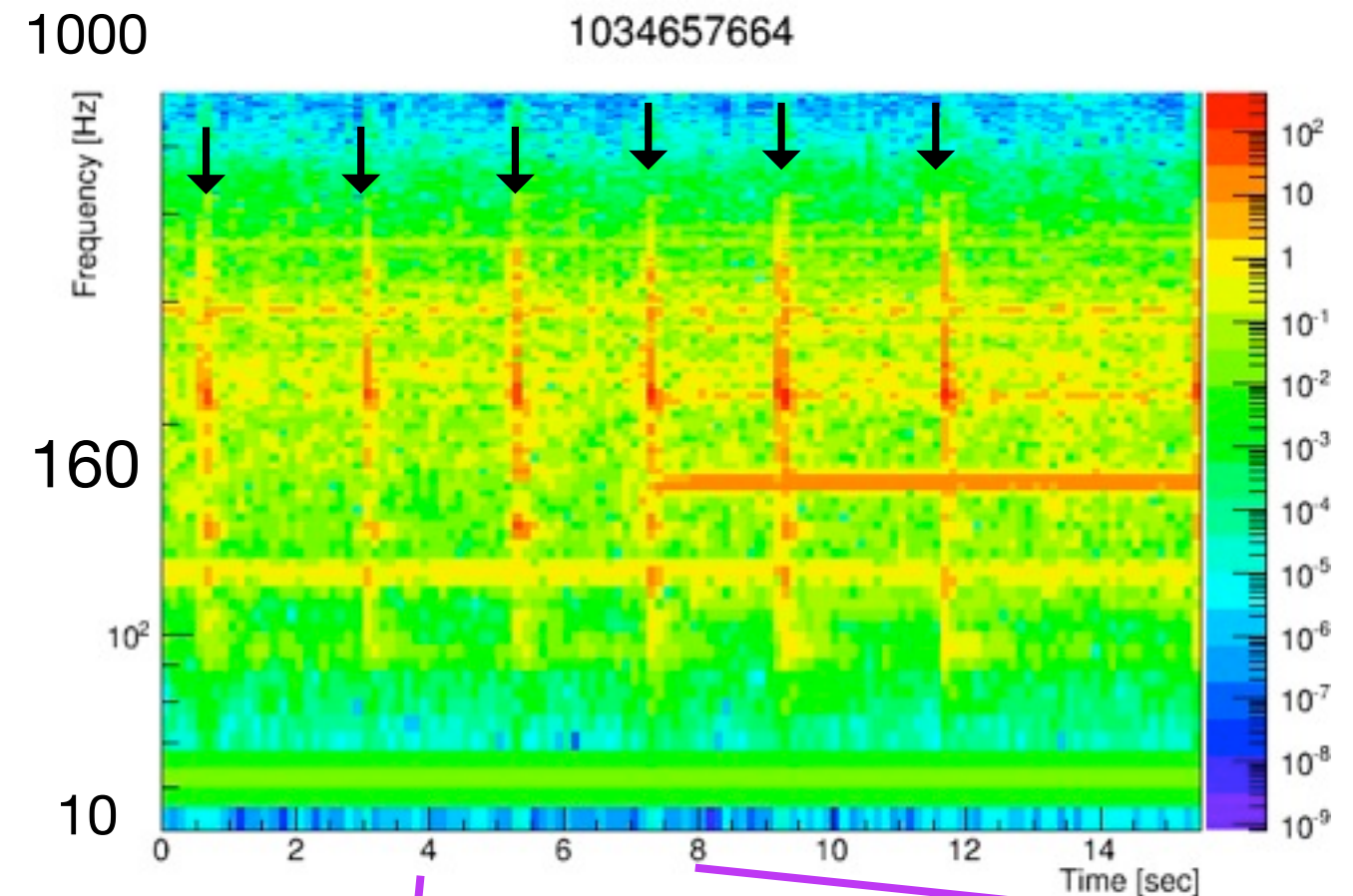
It is difficult to draw information of noise
⇒ spectrogram using Short-Fourier Transform

spectrogram of each channel data

GW channel

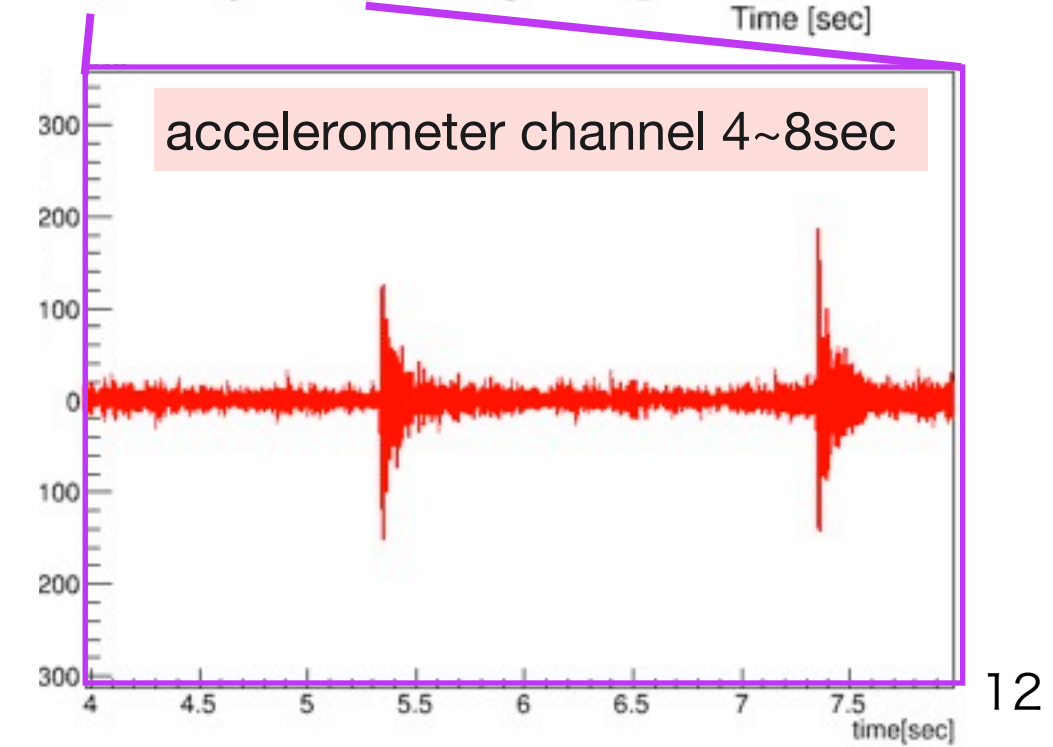


accelerometer channel



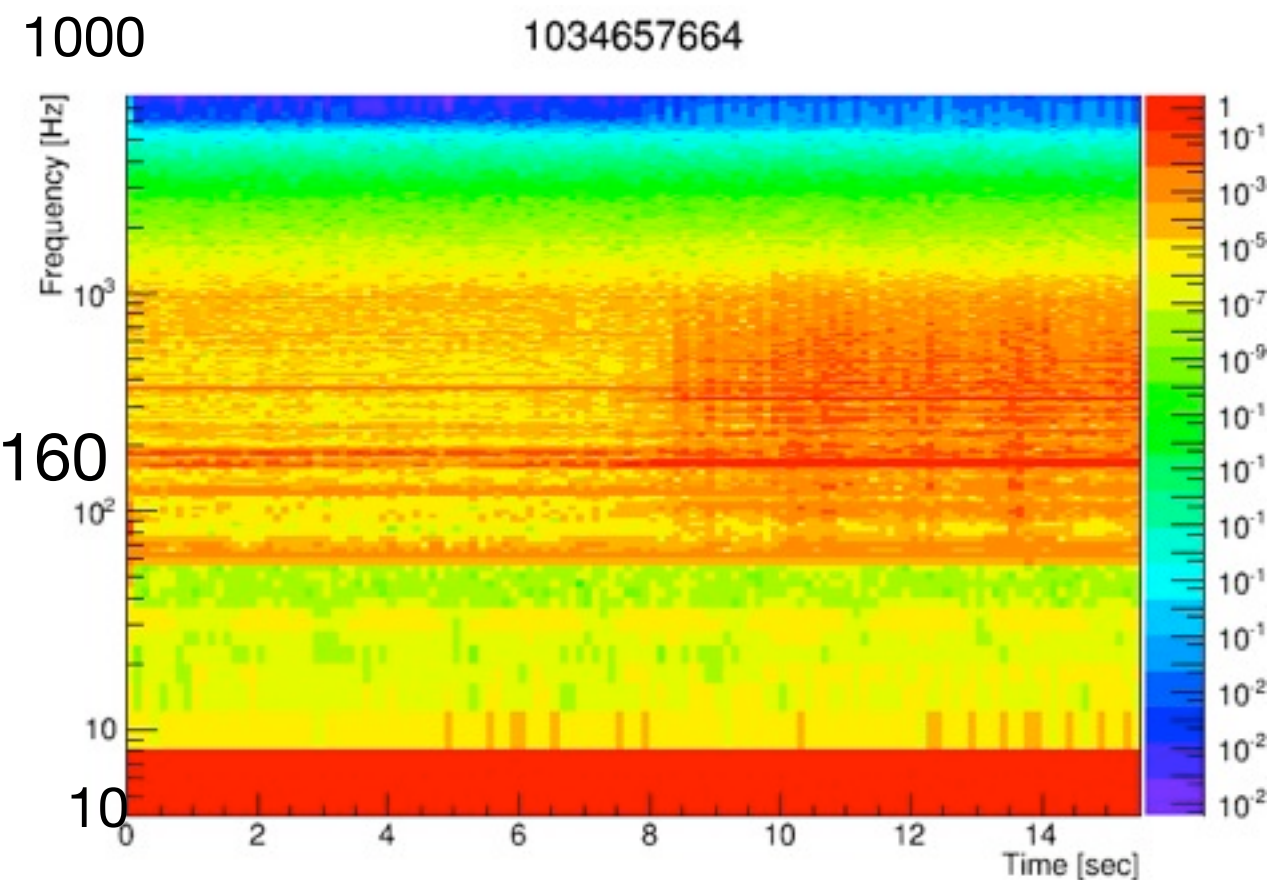
Black arrow in right figure show hardware injection equivalent (shock chamber directly)
⇒ There is this disturbance in GW channel?

We search correlation using developed correlation monitor.

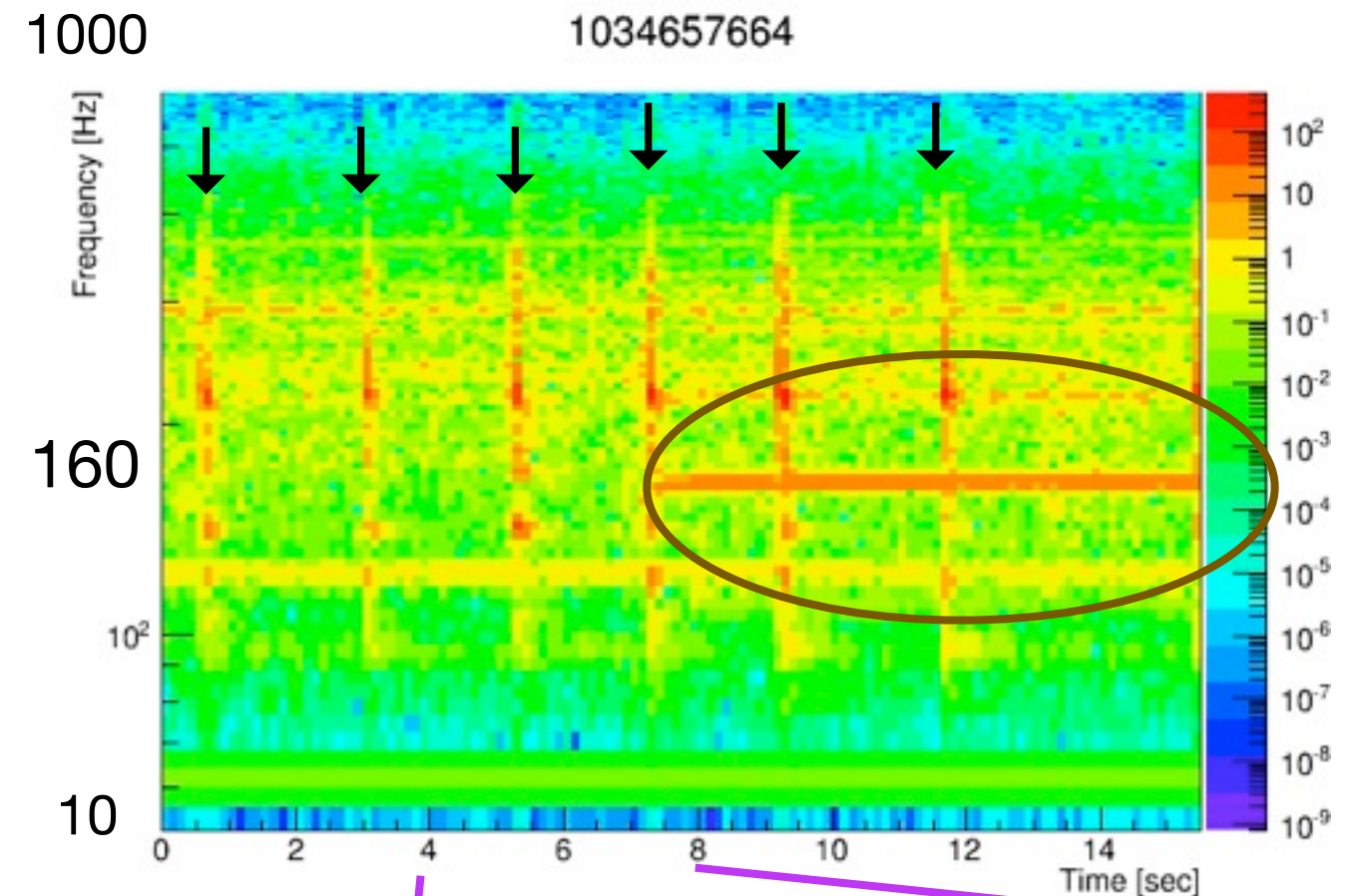


spectrogram of each channel data

GW channel



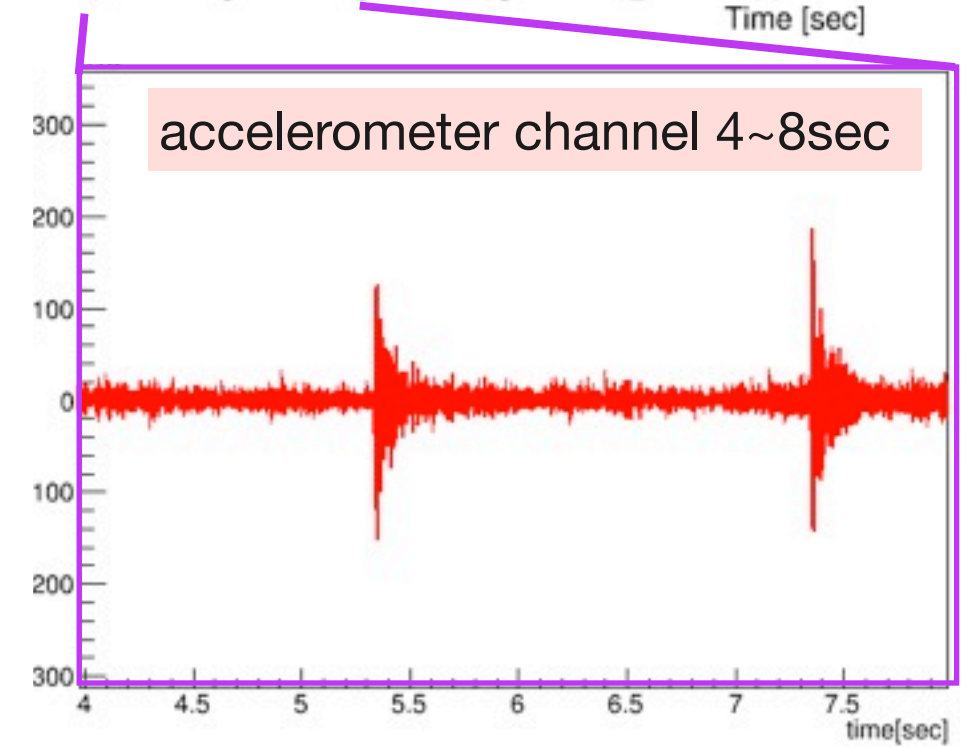
accelerometer channel



Black arrow in right figure show hardware injection equivalent (shock chamber directly)
⇒ There is this disturbance in GW channel?

We search correlation using developed correlation monitor.

accelerometer channel 4~8sec



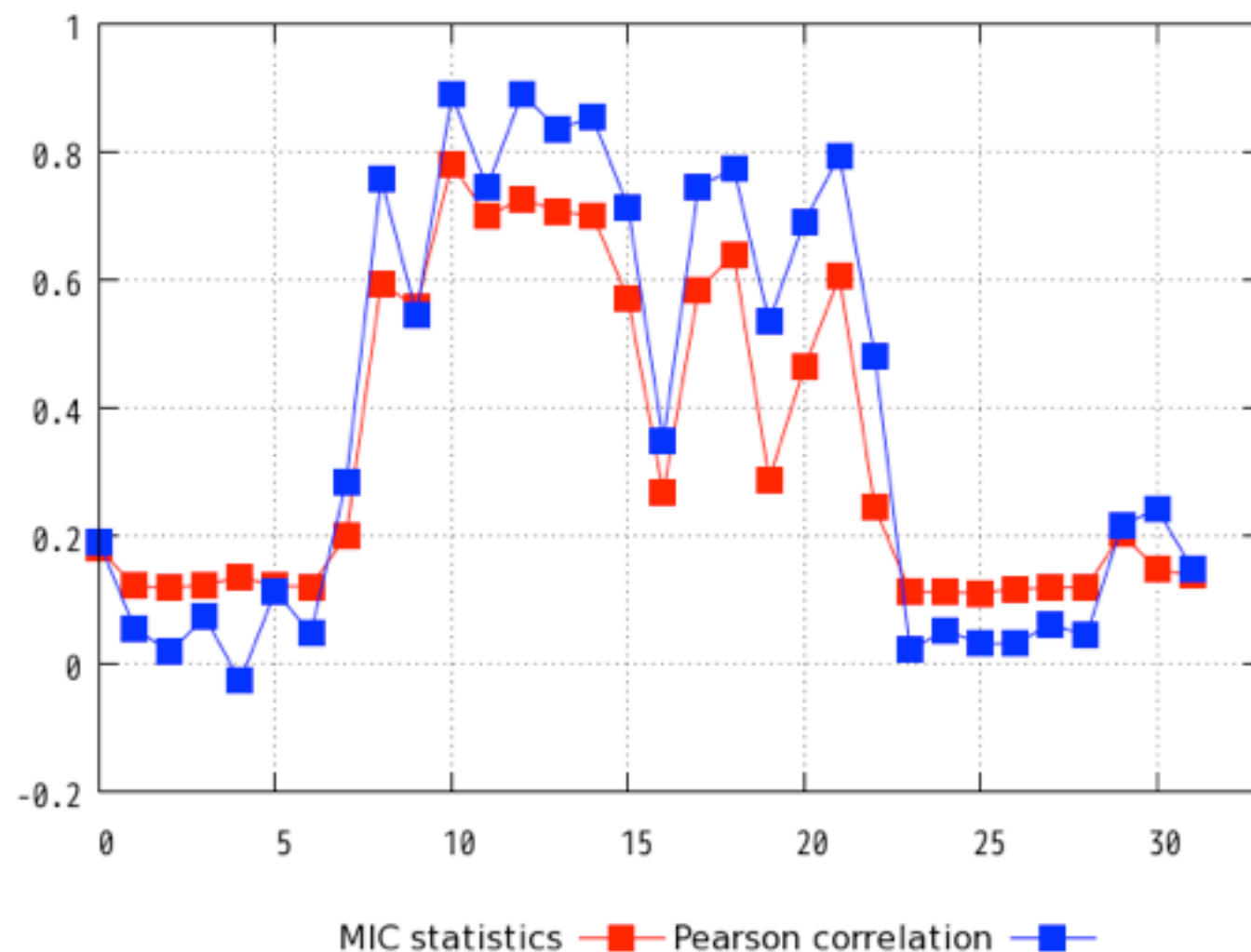
time change of correlation value

We search correlation using data set around hardware injection.

In spectrogram of accelerometer, the signal over 100Hz~200Hz

-> To pick up only this signal, data is band-pass filtered
with $f_{\text{low}}=100\text{Hz}$ and $f_{\text{high}} = 200\text{Hz}$

correlation value

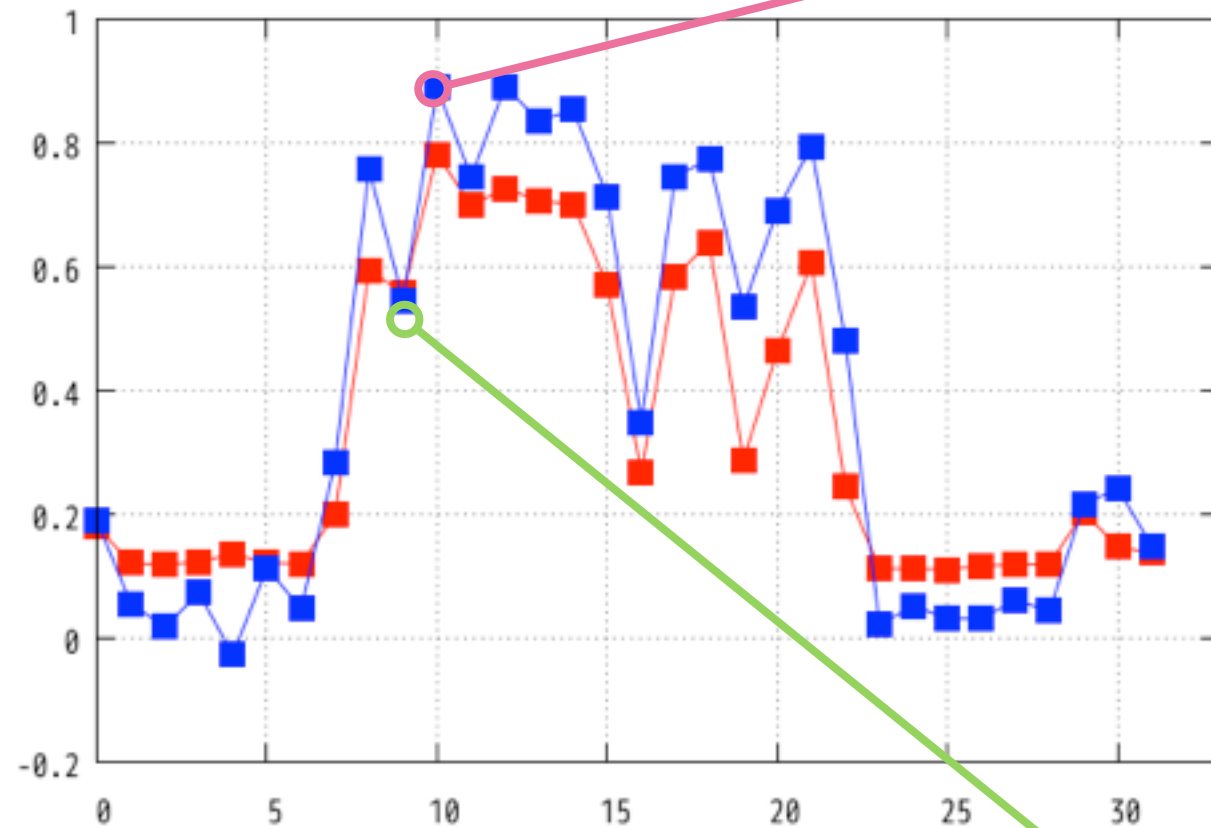


8sec ~ 21sec :
highly correlated value
with both Pearson and
MIC method

time domain data at each time

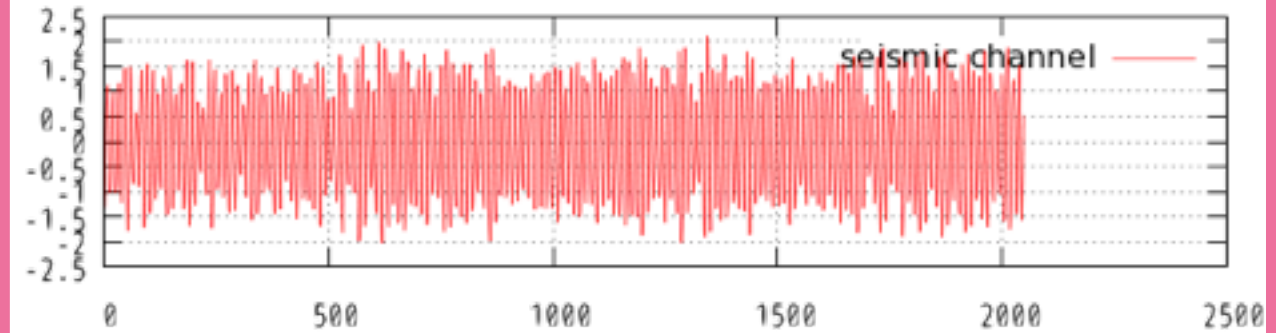
data is band-pass filtered with $f_{\text{low}}=100\text{Hz}$ and $f_{\text{high}} = 200\text{Hz}$ and normalized

correlation value

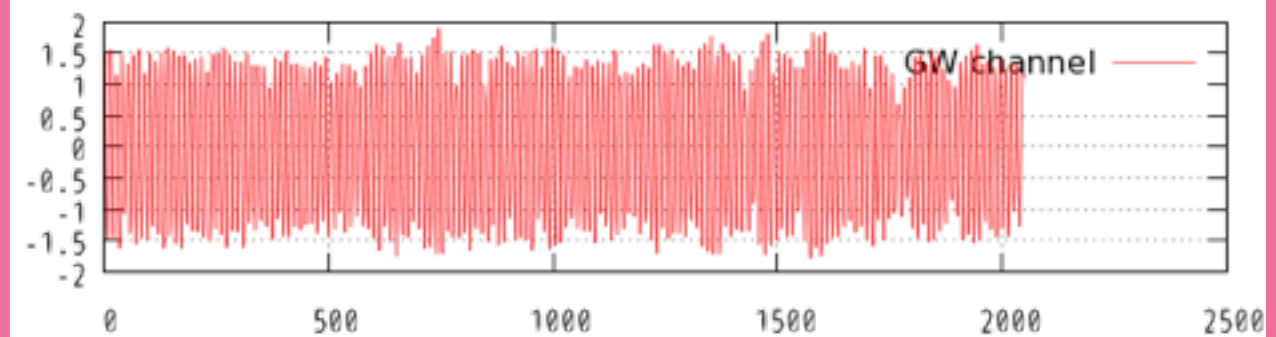


time [sec] MIC statistics — Pearson correlation —

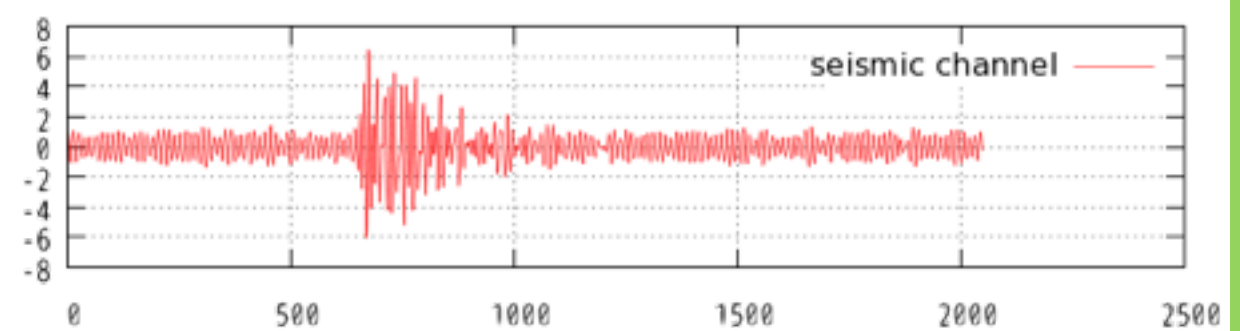
accelerometer channel rm timeshift = samples



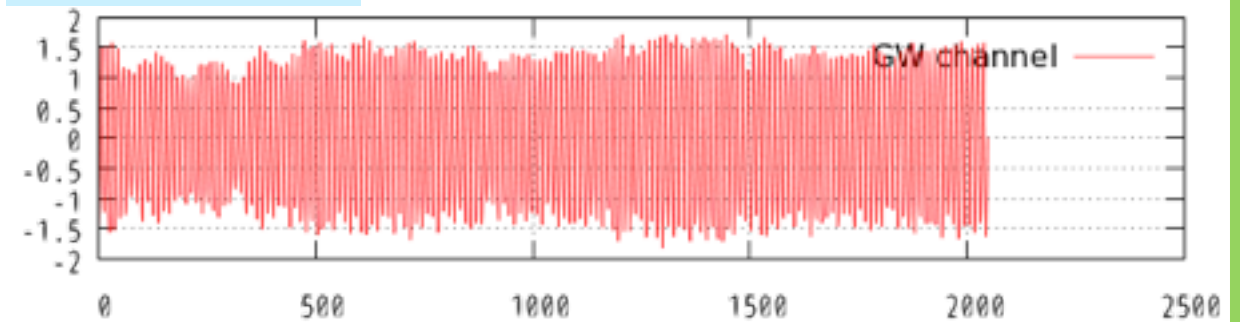
GW channel 57674_scatter_norm timeshift = samples



accelerometer channel m timeshift = samples



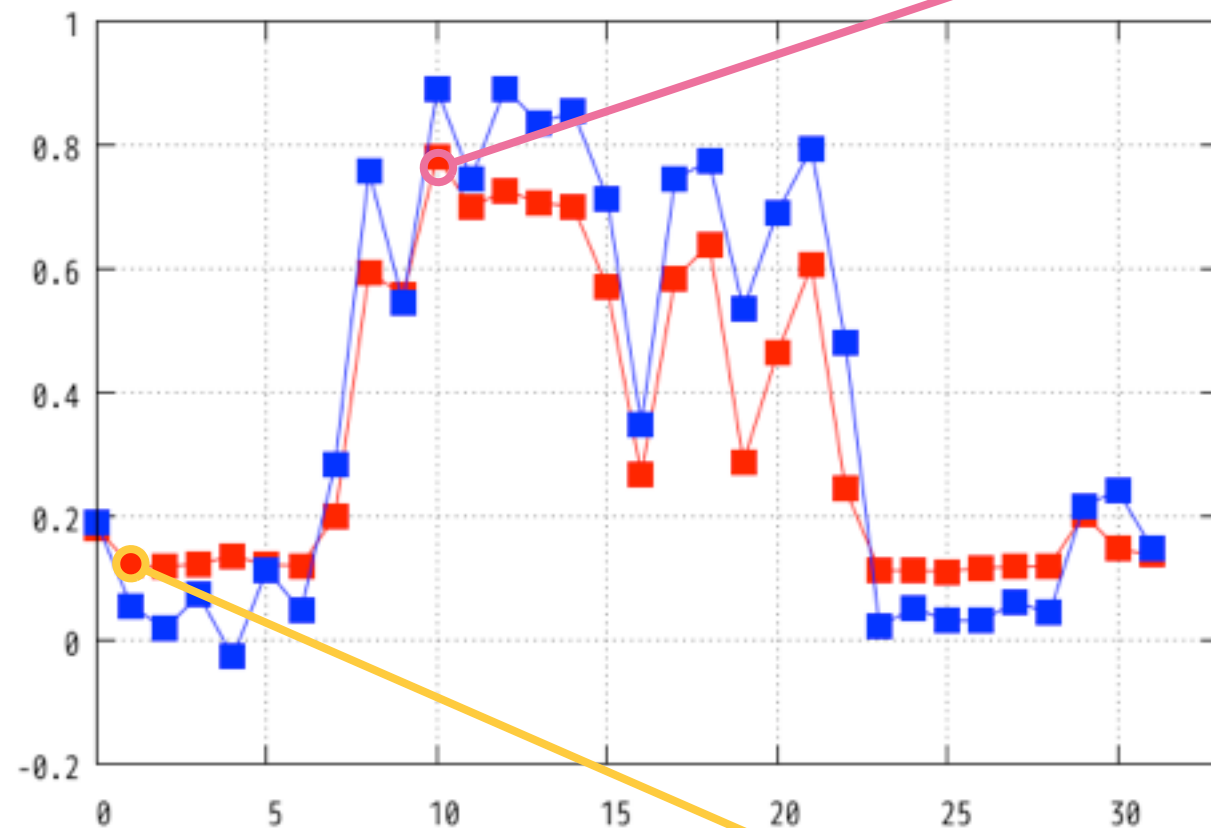
GW channel 57673_scatter_norm timeshift = samples



time change of scatter plot

data is band-pass filtered with $f_{\text{low}}=100\text{Hz}$ and $f_{\text{high}} = 200\text{Hz}$ and normalized

correlation value

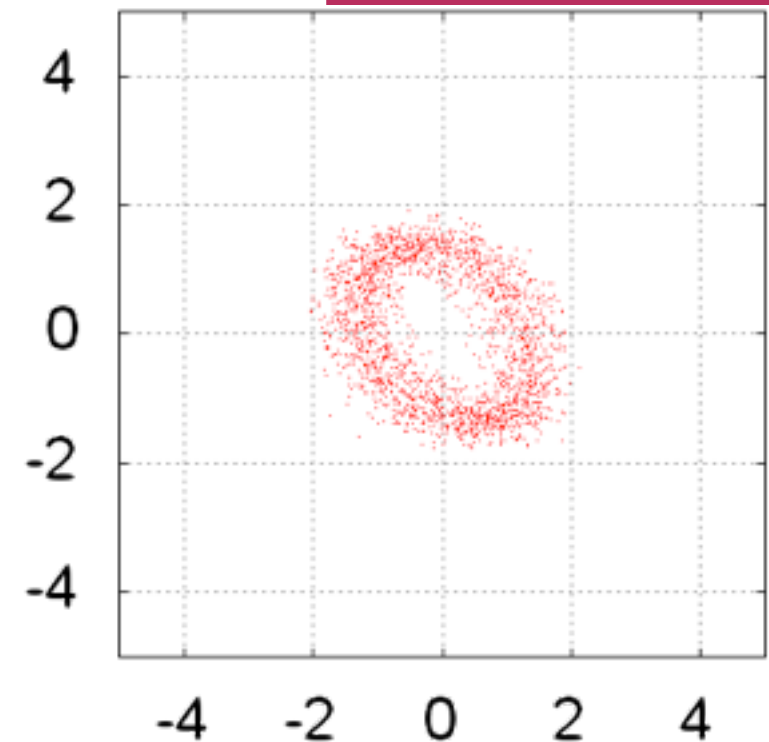


time [sec]

MIC statistics — Pearson correlation

correlated data set

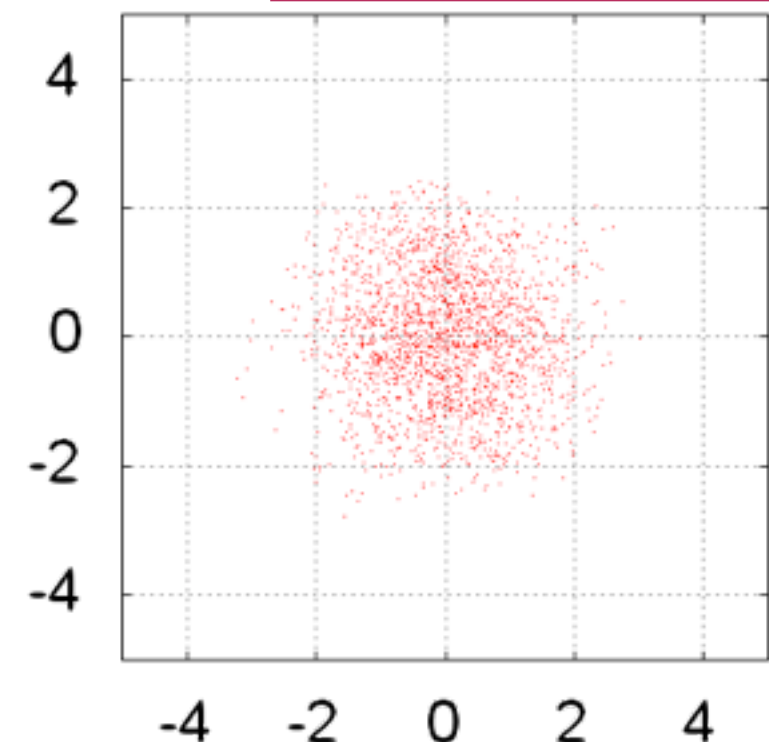
GW channel



accelerometer channel

uncorrelated data set

GW channel

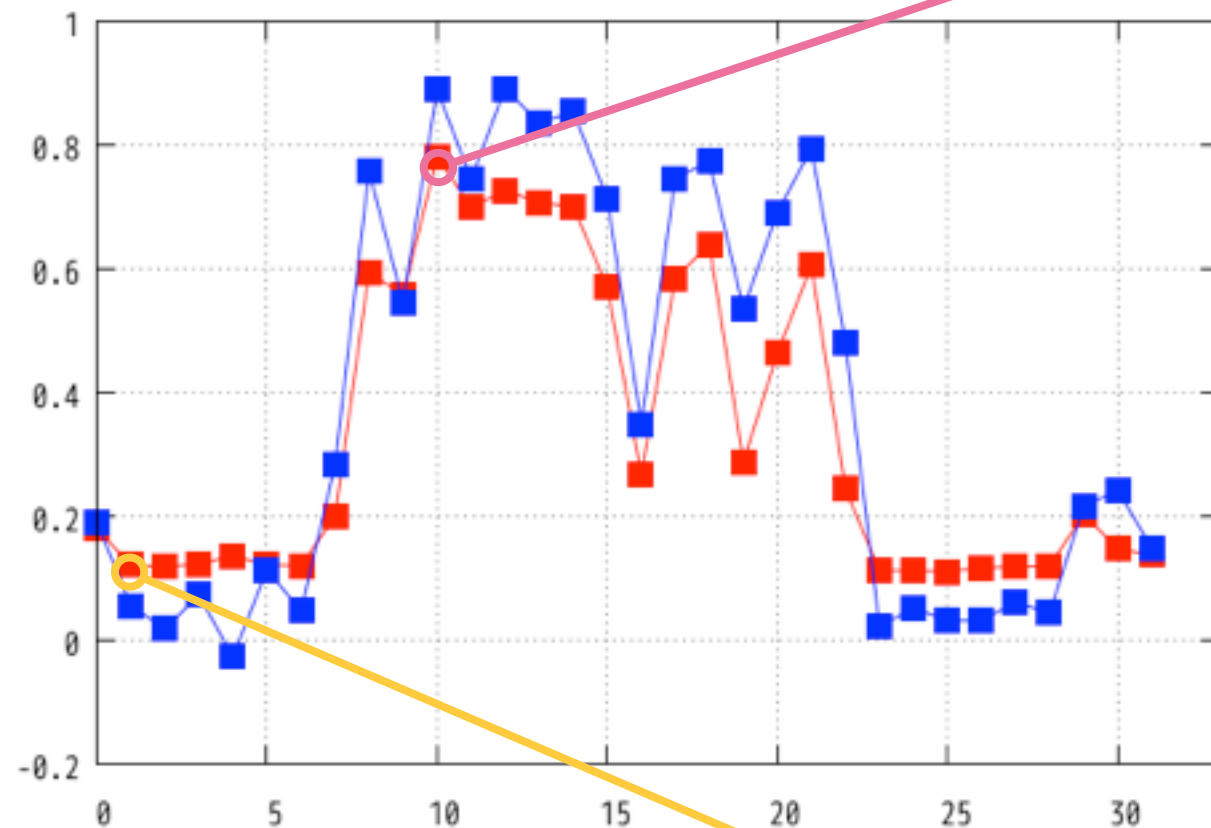


accelerometer channel

time change of scatter plot

data is band-pass filtered with $f_{\text{low}}=100\text{Hz}$ and $f_{\text{high}} = 200\text{Hz}$ and normalized

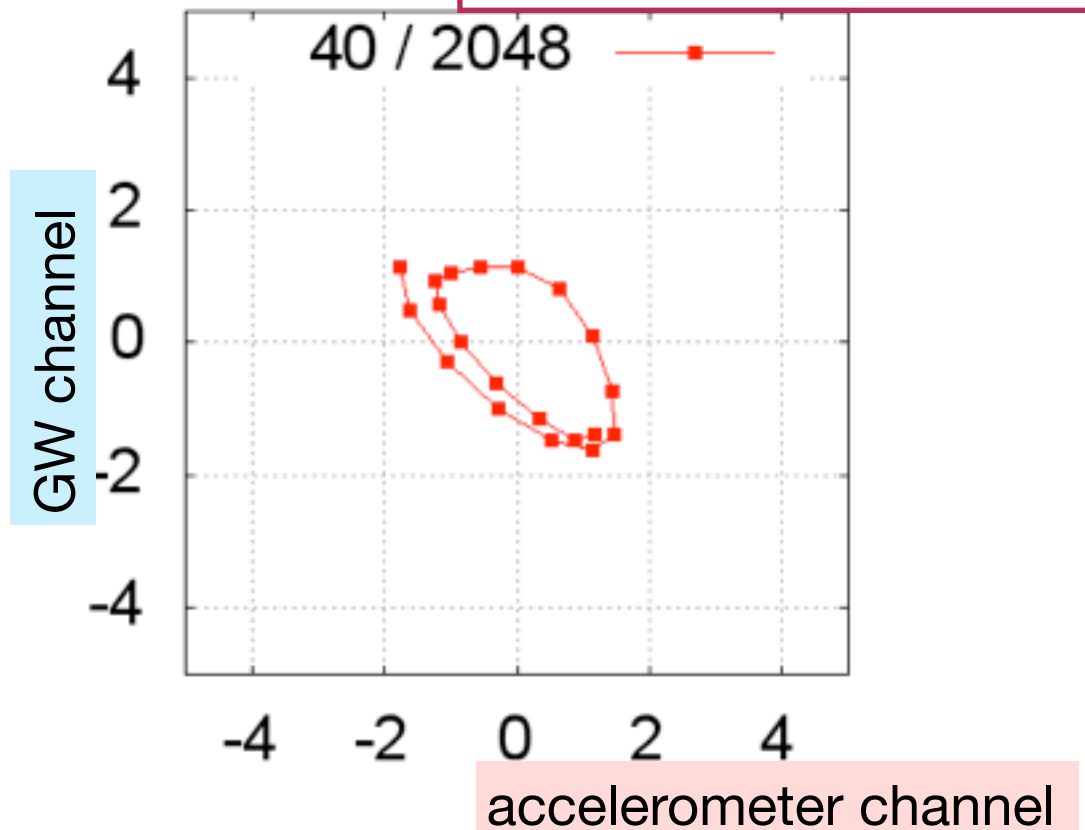
correlation value



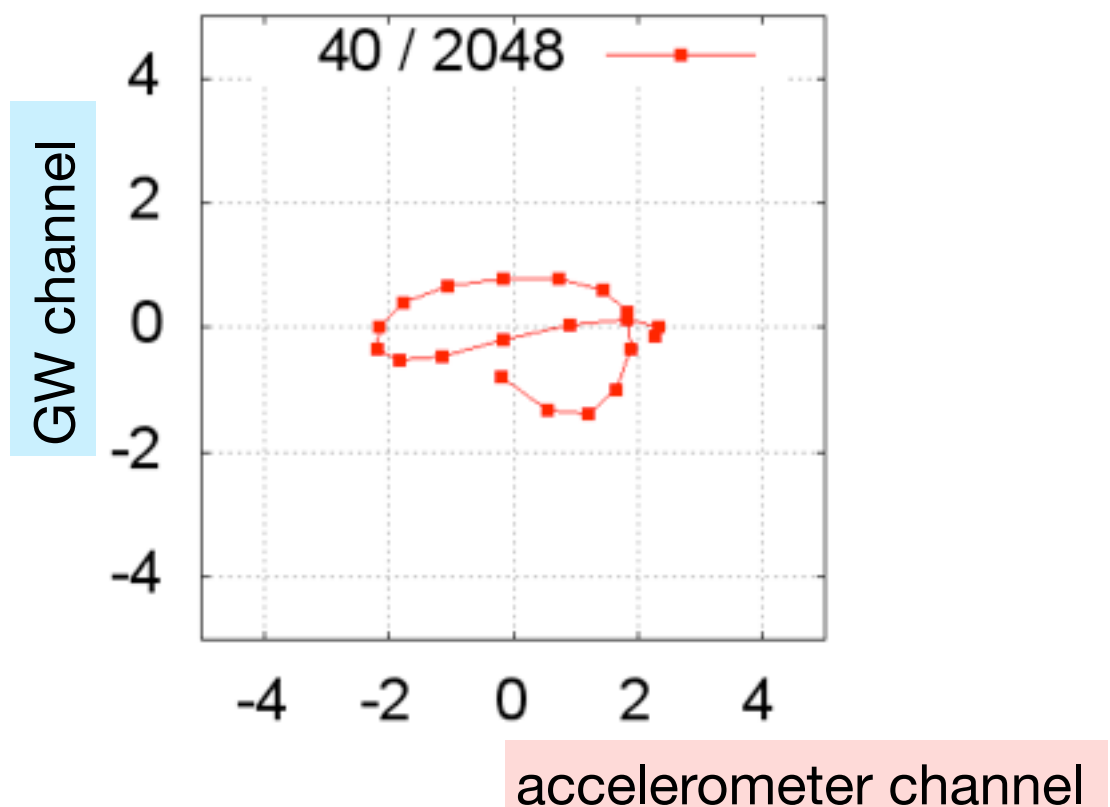
time [sec]

MIC statistics — Pearson correlation

correlated data set

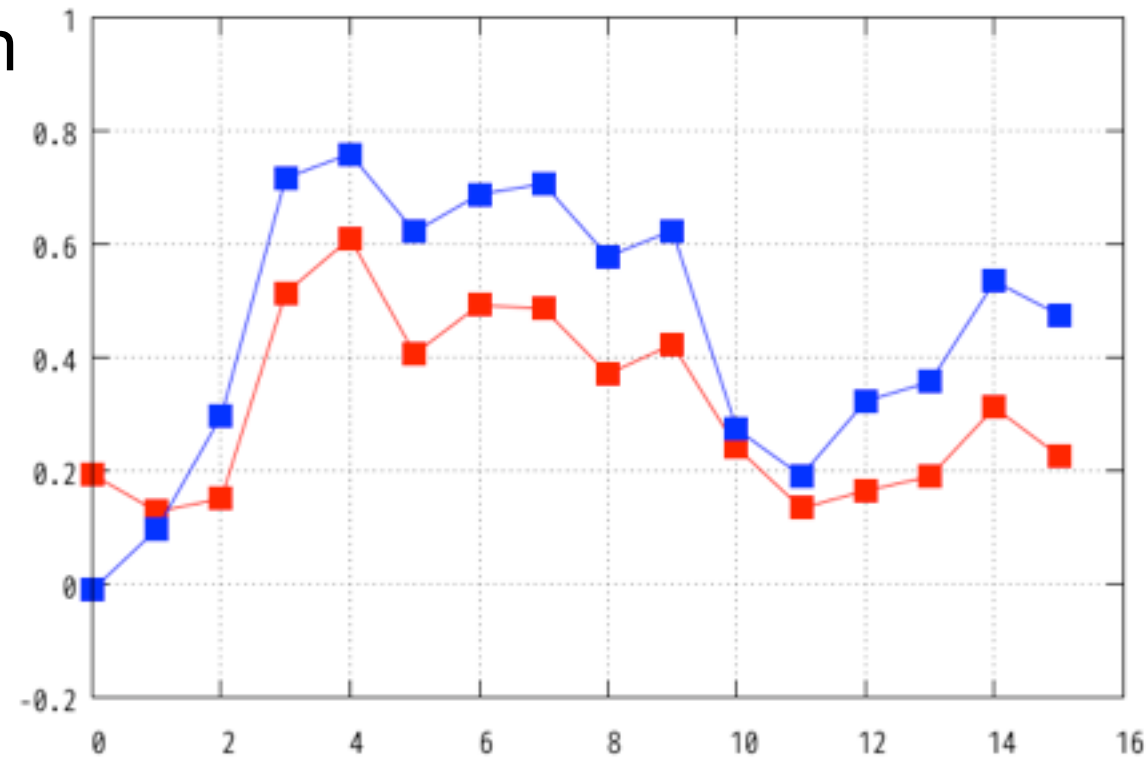


uncorrelated data set



We find same correlation in another data set

correlation
value

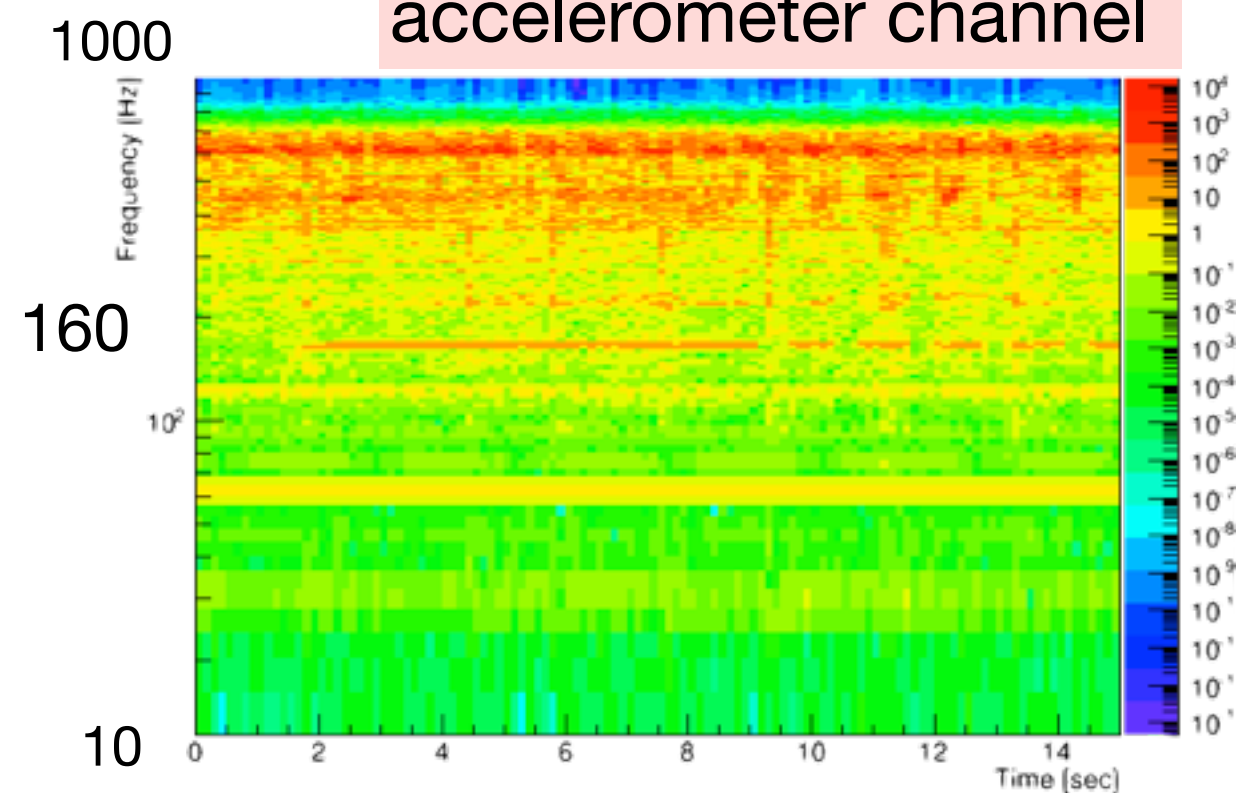
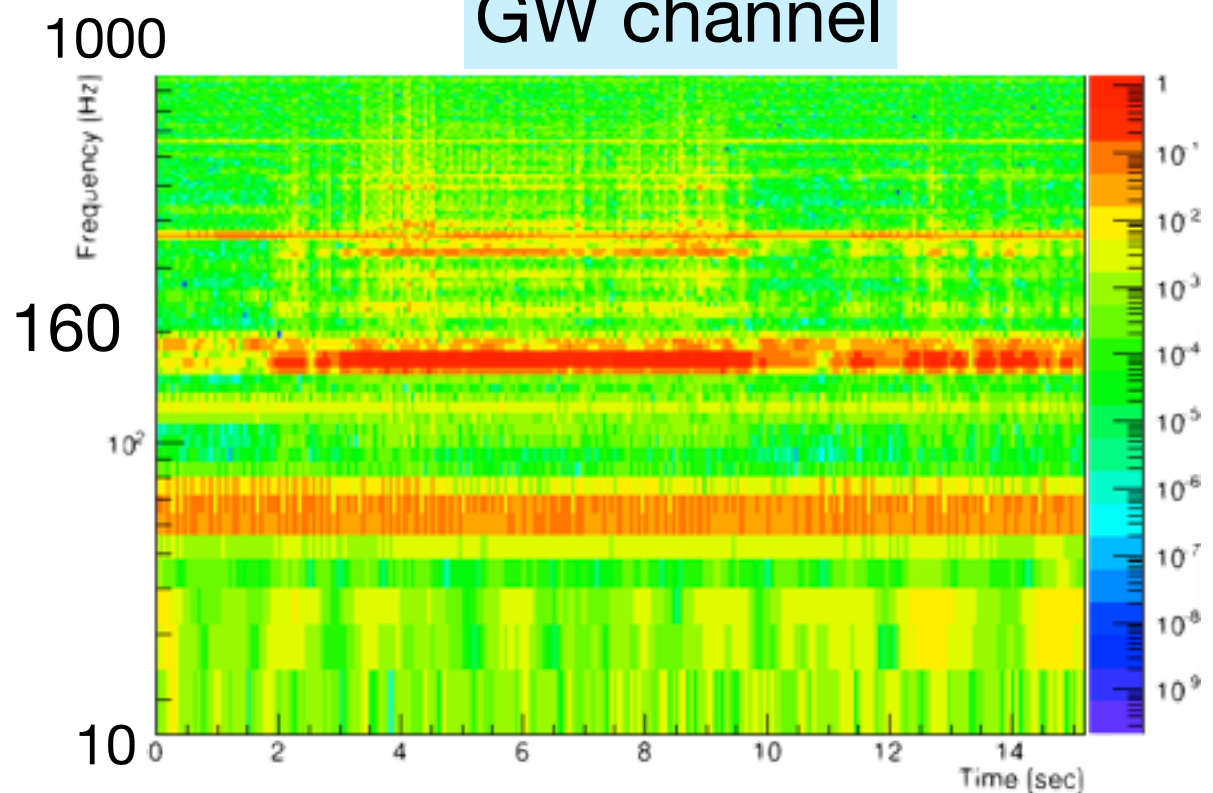


time [sec]

GW channel

MIC statistics — Pearson correlation

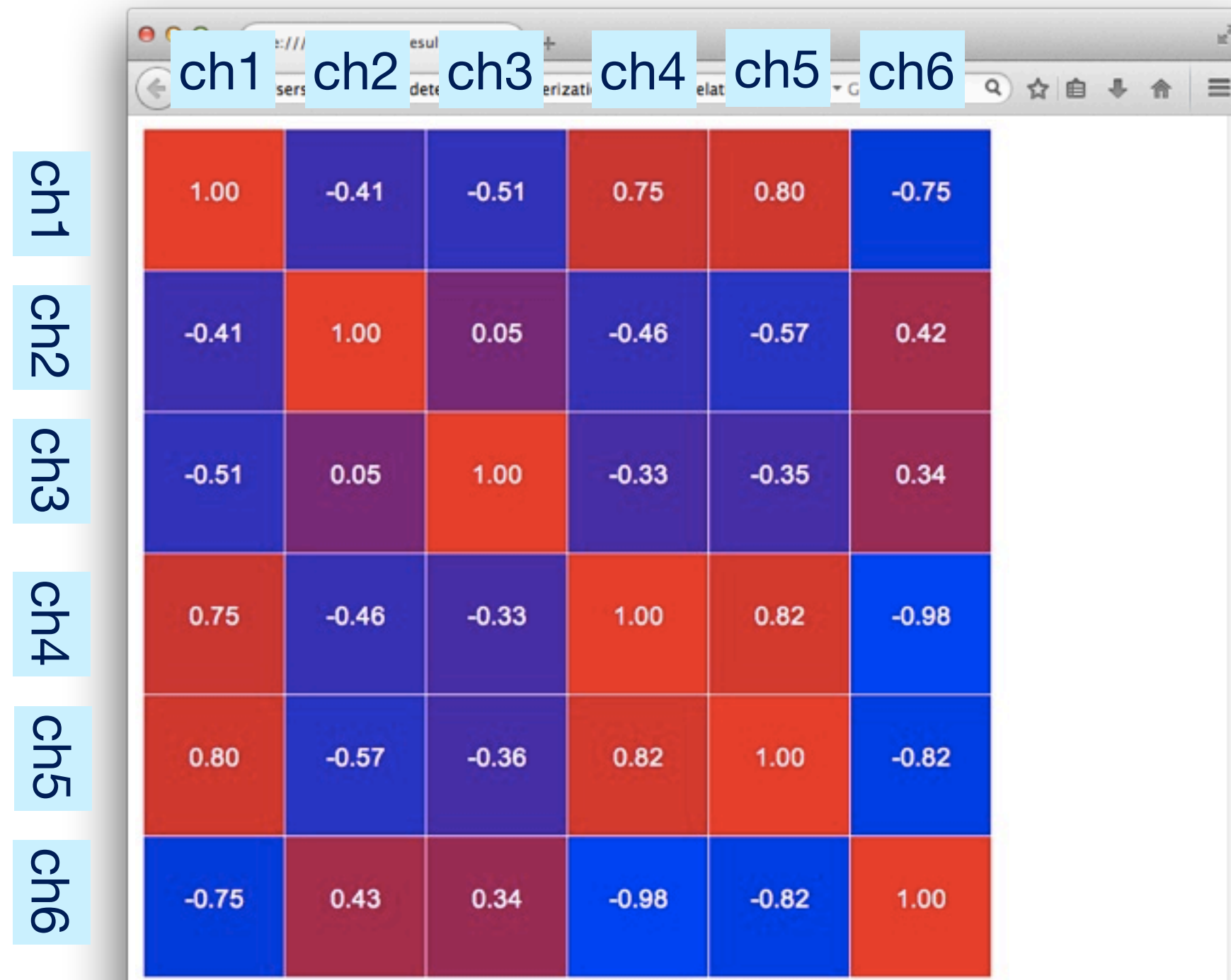
accelerometer channel



correlation map (very preliminary)

In this study, only two channel data is used. But in future, number of channel will increase.

So viewer page was also developed to observe easily which channels are correlated.



Note : Used data is meaningless.

Summary

- Many linear and non-linear correlated noises are observed in LIGO and Virgo. These noise limit detector sensitivity. Correlation analysis in this study will localize the noise source. As a result, false triggers generated GW search pipeline can be removed and increase detection efficiency.
- In KAGRA's commissioning and observation phase, localization of same correlated noise is important. In order to detect both linear and non-linear correlation, we developed monitor tool using
 - Pearson correlation coefficient**
 - Maximum Information Coefficient(MIC).**The monitor tool which cover non-linear correlation is **world-first**.
- Correlation analysis using GW channel and accelerometer of CLIO detector data. Correlation monitor tool discover a correlation in two channels.

