

# Interferometer design of LCGT



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# 干渉計 再検討事項

Two nearby interferometers

As independent as possible

Center rooms: separated

3km vacuum tube: common

## Main interferometer

Baseline Length : 3km

Broad band RSE

with power recycling

Arm cavity Finesse : 1550

Power Recycling Gain : 11

Signal band Gain : 15

Stored Power : 771kW

Signal band : 230Hz



Single interferometer

New IFO layout

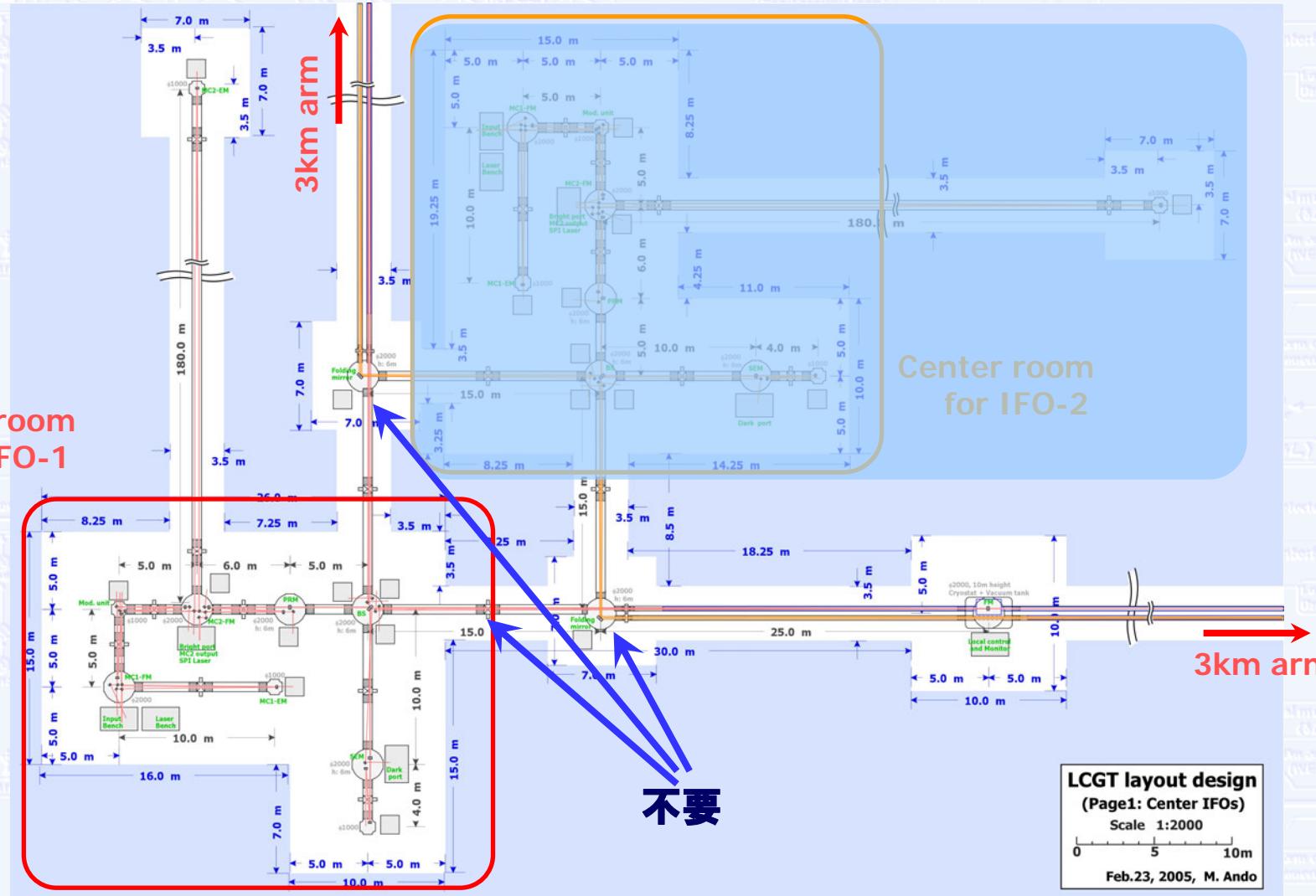


Baseline Length : 2.7km

Parameter optimization

Observable range

# Interferometer layout



# Optical readout noise

LCGT sensitivity : Mostly limited by quantum noises  
(Radiation pressure noise and Shot noise)

## Shot noise

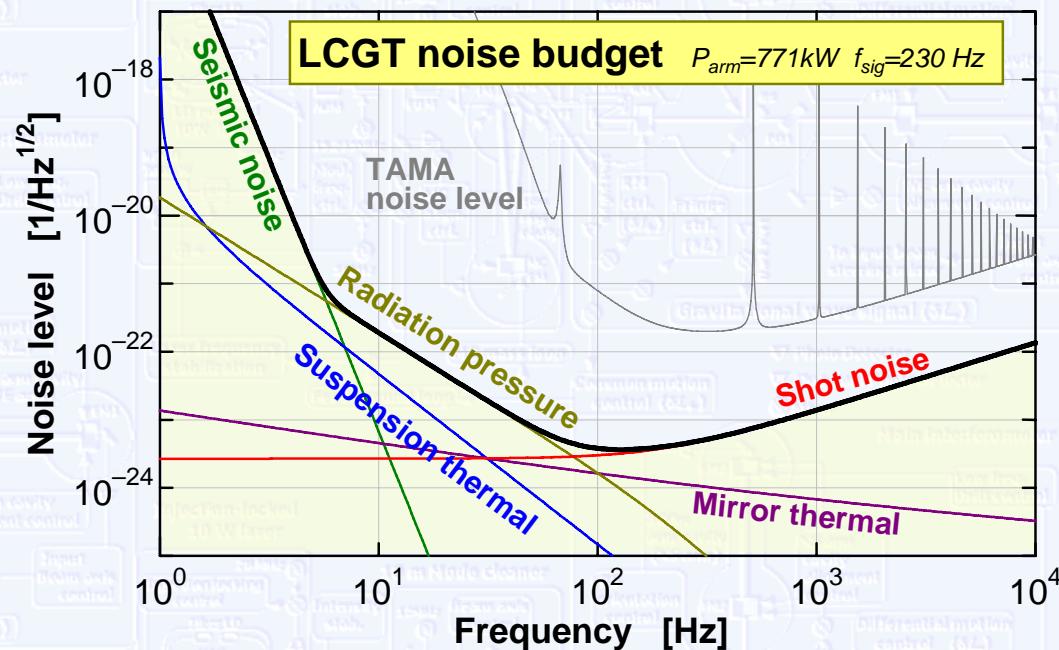
Phase fluctuation of light  
Proportional to  $P^{-1/2}$

## Radiation pressure noise

Amplitude fluctuation of light  
Proportional to  $P^{1/2}$



Noise level depends on  
Laser power in IFO  
Signal band of IFO



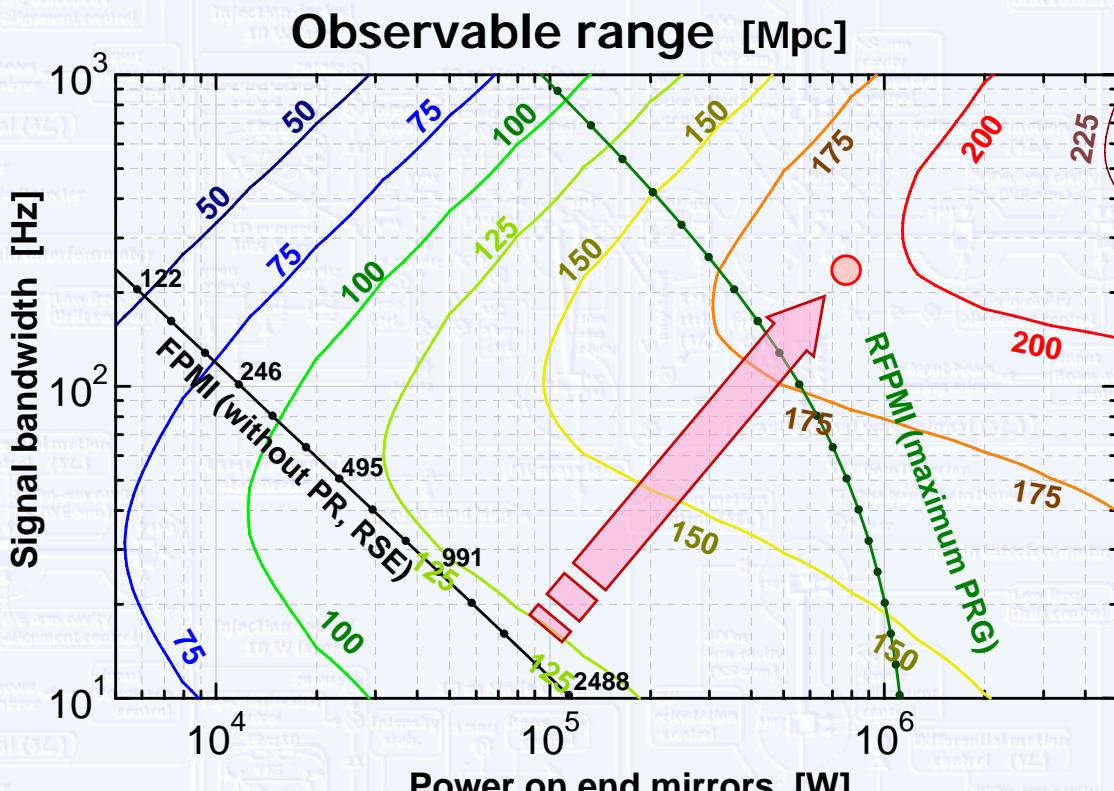
Design the interferometer optical parameters

# Optical parameter selection

Optimize for  $1.4M_{\text{solar}}$  NS inspiral  
Realistic parameters

Calculate observable range  
as a function of  
Light power in cavities  
Signal bandwidth

Arm cavity Finesse : 1550  
Power Recycling Gain : 11  
Signal Band Gain : 15  
Stored Power : 771kW  
Signal band : 230Hz



Observable range : 185Mpc

# Optical parameter selection

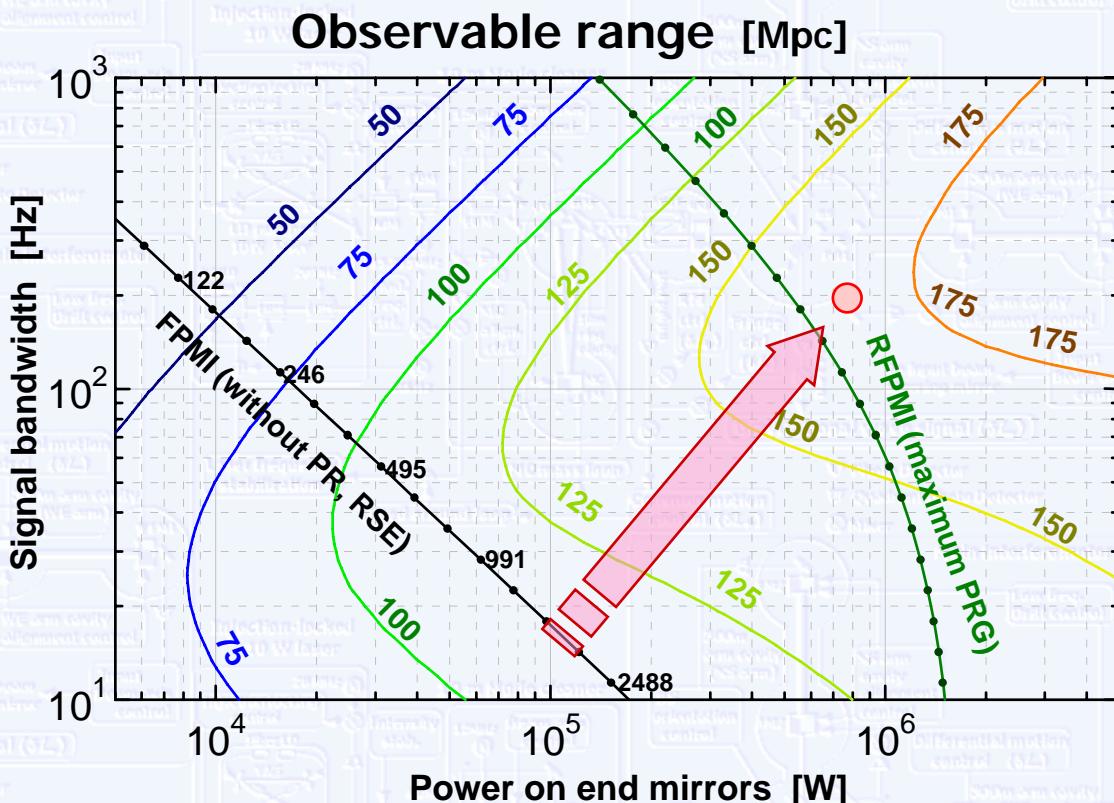
Baseline length → 2.7km

Optimize for  $1.4M_{\text{solar}}$  NS inspiral  
Realistic parameters

Calculate observable range  
as a function of  
Light power in cavities  
Signal bandwidth

Arm cavity Finesse : 1550  
Power Recycling Gain : 11  
Signal Band Gain : 11  
Stored Power : 771kW  
Signal band : 195Hz

(New design)



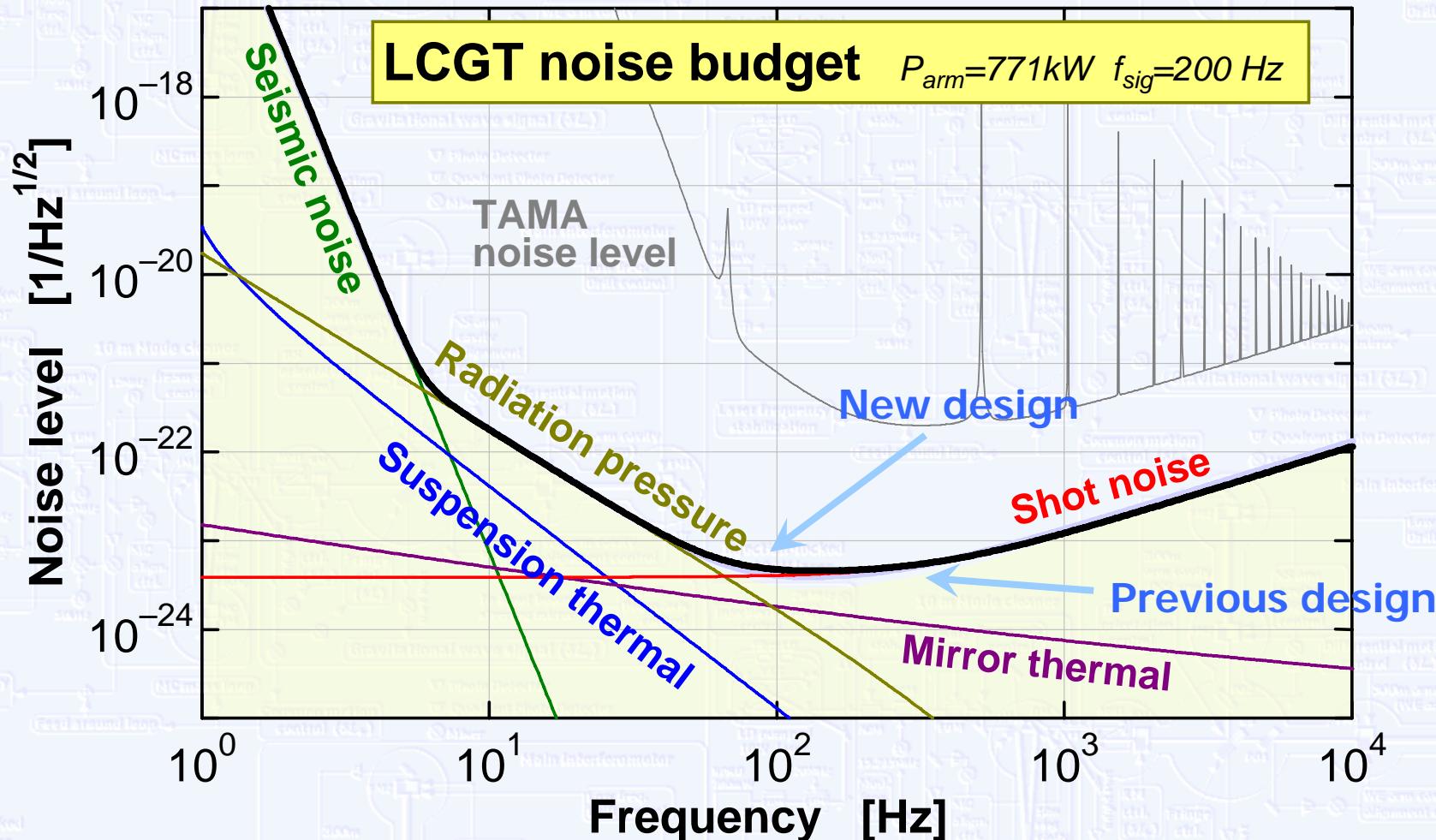
Observable range : 167Mpc  
(10% reduction,  
→ event rate ~25% reduction)

SNR = 10  
Single interferometer  
Optimal direction and polarization

# Sensitivity Curve

Almost no difference at a glance

$$3.95 \times 10^{-24} \text{ Hz}^{-1/2} \rightarrow 4.58 \times 10^{-24} \text{ Hz}^{-1/2}$$



# 感度向上の可能性

## ハイパワー化

レーザー光源 (出力 150W)

入出射光学系 (透過率 50%)

鏡の品質 (光損失 10ppm)

パワーリサイクリング (ゲイン 11)

→ パワー 2倍

## 狭帯域化

Detuned RSE

## スカイージング

ダークポートからの入射 ~10dB?

## 鏡の大型化

現状: 直径 25cm, 厚さ 15cm, 30kg

→ 直径 30cm, 42kg

7%程度のレンジ向上  
(ただし鏡冷却 → FM透過 3kWが限界  
現状 0.39kW)

20%程度のレンジ向上  
(70-80%程度のイベントレート向上)  
が期待できる

70%程度のレンジ向上  
(5倍程度のイベントレート向上)  
が期待できる

13%程度のレンジ向上  
(40%程度のイベントレート向上)  
が期待できる

# 感度向上の可能性

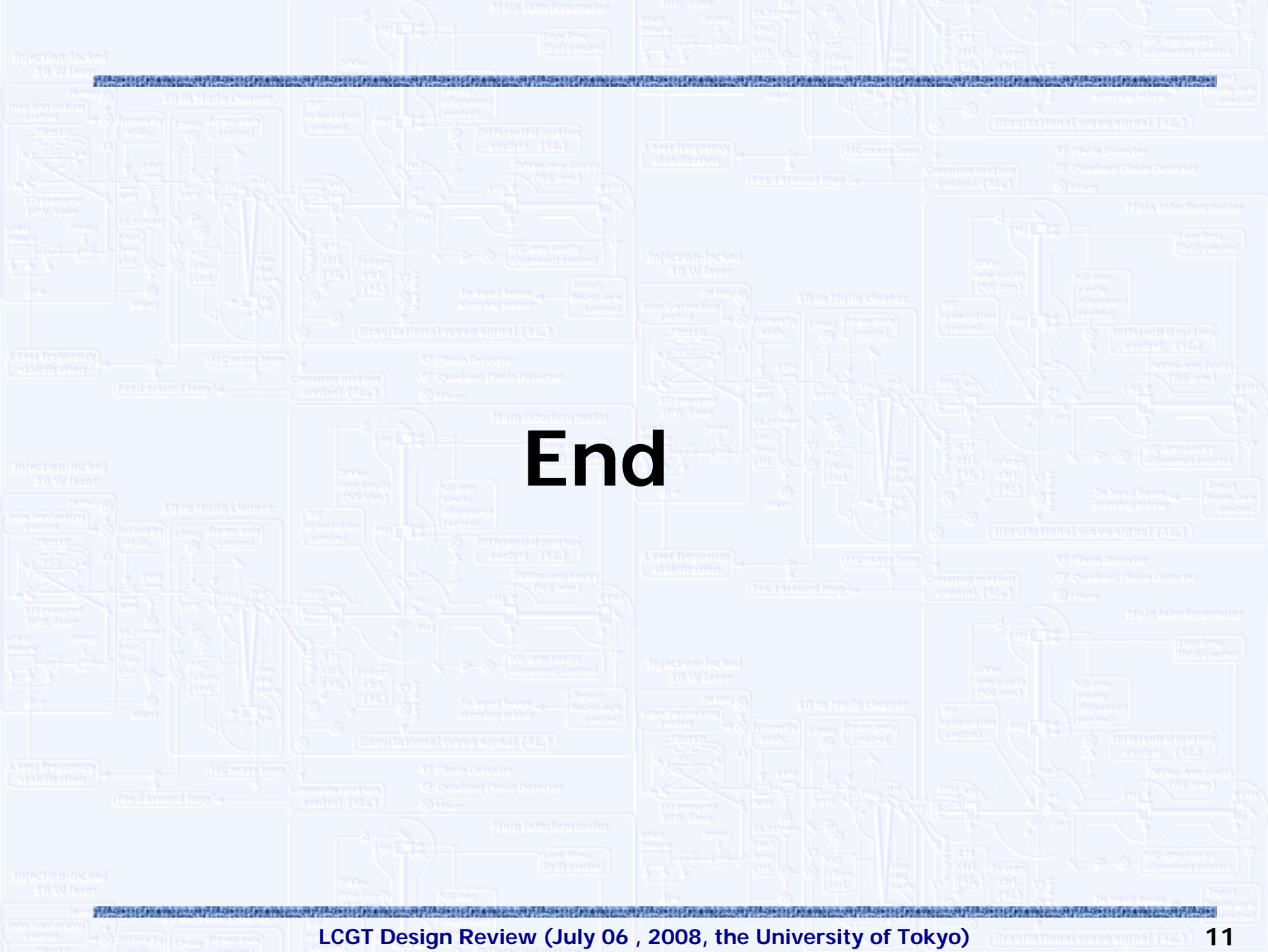
	Obs. range [Mpc]	Range gain	Rate ratio
Original design Baseline 3km	185 (103)	10%	1.33
Current design Baseline 2.7km	167 ( 93)	---	1
<b>High Laser Power</b> Power in arm x2	179 (100)	7%	1.23
<b>Narrow band</b> Detuned RSE	200 (112)	20%	1.72
<b>Squeezing</b> Shot noise -10dB	284 (159)	70%	4.92
<b>Heavy mirror</b> 30kg → 42kg	190 (106)	13%	1.47
Adv. LIGO	320 (180)	92%	7.03
Adv. VIRGO	215 (120)	29%	2.13

# まとめ

基線長が 3km → 2.7kmに短縮されたことにより  
観測レンジは 184Mpc → 167Mpc に低下する  
イベント数は 約3/4になる

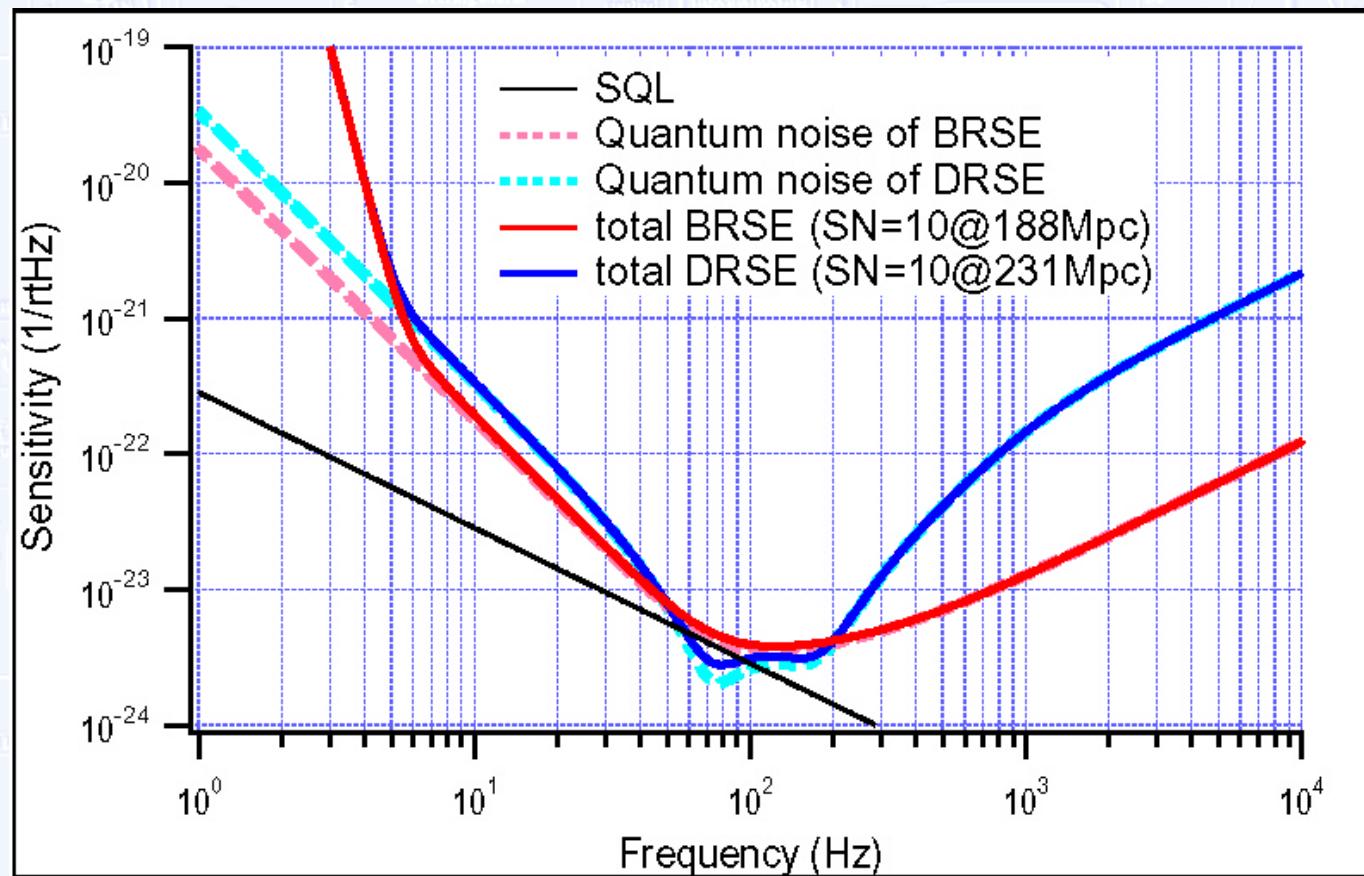
提言： LCGT もしくは Ad. LCGT として、  
感度向上の方策を持っているべき

# End



# Broadband - Narrowband

Detuned RSE configuration  
→ SNR increase 23%  
Event rate increase 86%



# Comparison of detectors

## • Comparison of next generation GW detectors

### LCGT (JPN)

2 detectors (3km)  
(2 nearby detectors)

Long baseline  
Better seismic  
attenuation system  
Underground site

Low-mechanical-loss  
mirrors and suspensions  
Cryogenic (20k)

High-power laser source  
Low-loss optics  
Broad-band RSE config.

### Scale

Seismic noise  
reduction

Thermal noise  
reduction

Quantum noise  
reduction

### Advanced LIGO (USA)

3 detectors (4km)  
(2 nearby, 1 separated)

Long baseline  
Better seismic  
attenuation system  
Suburban site

Low-mechanical-loss  
mirrors and suspensions  
Flat-top beam

High-power laser source  
Low-loss optics  
Detuned RSE config.

# LCGT optical configuration

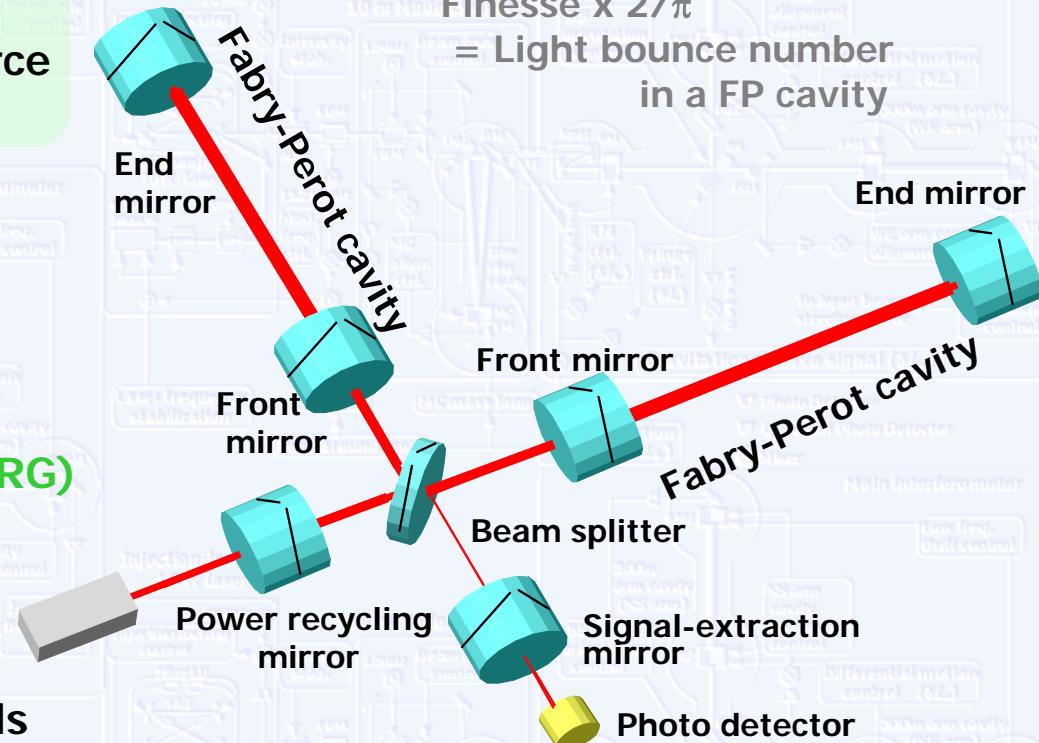
## Resonant-sideband extraction with power recycling

High-finesse arm cavities  
PRM between BS and laser source  
SEM at the detection port

Power recycling  
PRM+FM  
→ Increase effective finesse  
→ Increase power in cavities  
by Power-recycling gain (PRG)

Resonant-sideband extraction  
SEM+FM  
→ Decrease  
effective finesse for signals  
→ Increase signal band  
by Signal-band gain (SBG)

Finesse  $\times 2/\pi$   
= Light bounce number  
in a FP cavity



# Interferometer overview

## Two nearby interferometers

As independent as possible  
Center rooms: separated  
3km vacuum tube: common

## Main interferometer

Broad band RSE  
with power recycling  
Arm cavity Finesse : 1550  
Power Recycling Gain : 11  
Signal band Gain : 15  
Stored Power : 771kW  
Signal band : 230Hz

## Input optics

Two mode cleaners  
Baseline lengths : 10m and 180m  
Modulators for IFO control  
Laser stabilization

## Output optics

Output mode cleaner  
Round trip length : 70mm  
Multiple InGaAs photo diodes

## Observation system

Monitor and organize  
the whole detector system  
Automatic lock-acquisition  
Automatic interferometer adjustment  
Monitor and diagnosis

(Previous design)

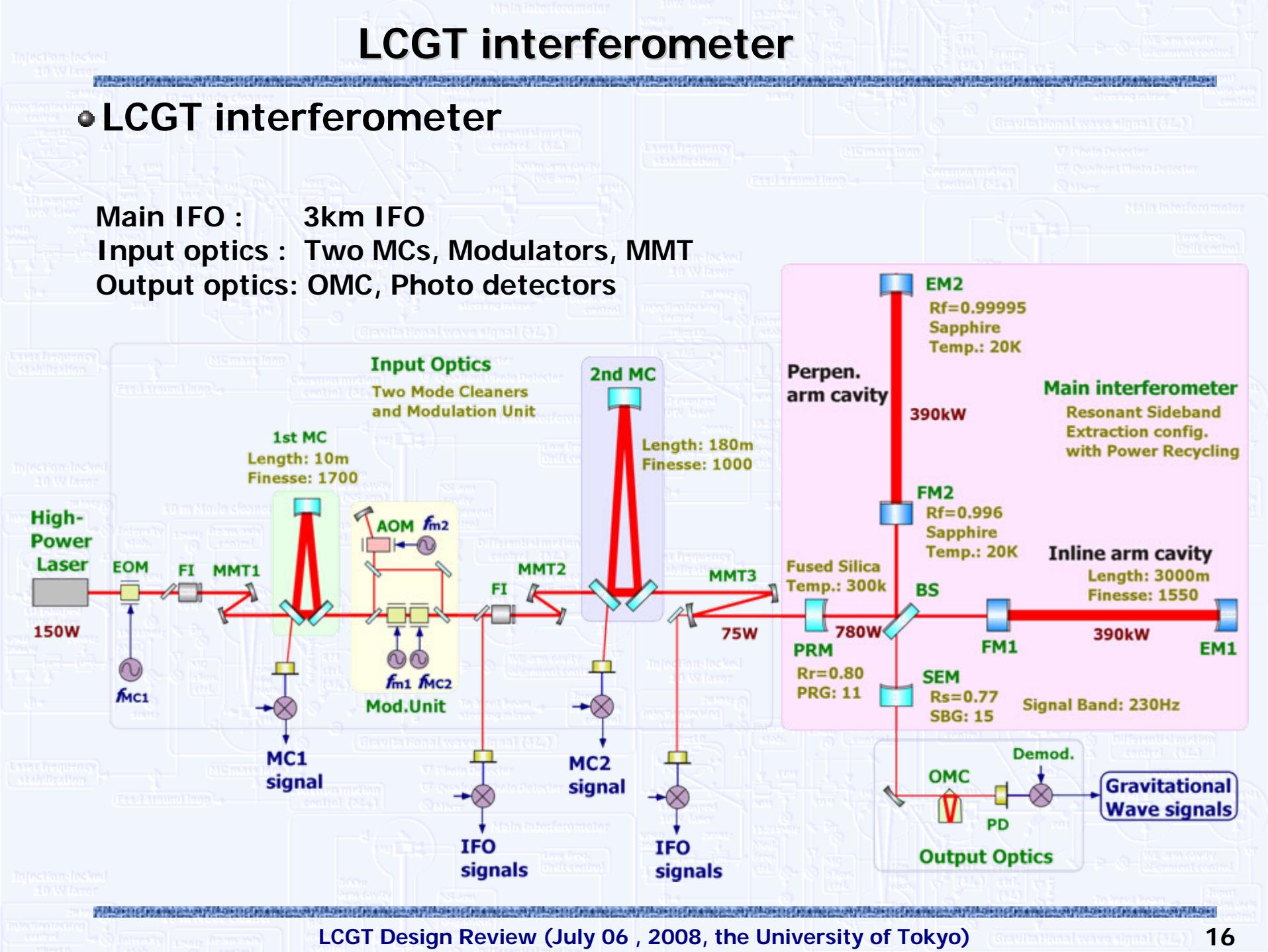
# LCGT interferometer

## LCGT interferometer

Main IFO : 3km IFO

Input optics : Two MCs, Modulators, MMT

Output optics: OMC, Photo detectors



# Main interferometer (2)

## • Interferometer optical configuration

Ideally, possible to realize  
same power and signal BW  
with any config.

Power : cavity finesse, PRM

Signal BW : cavity finesse, SRM

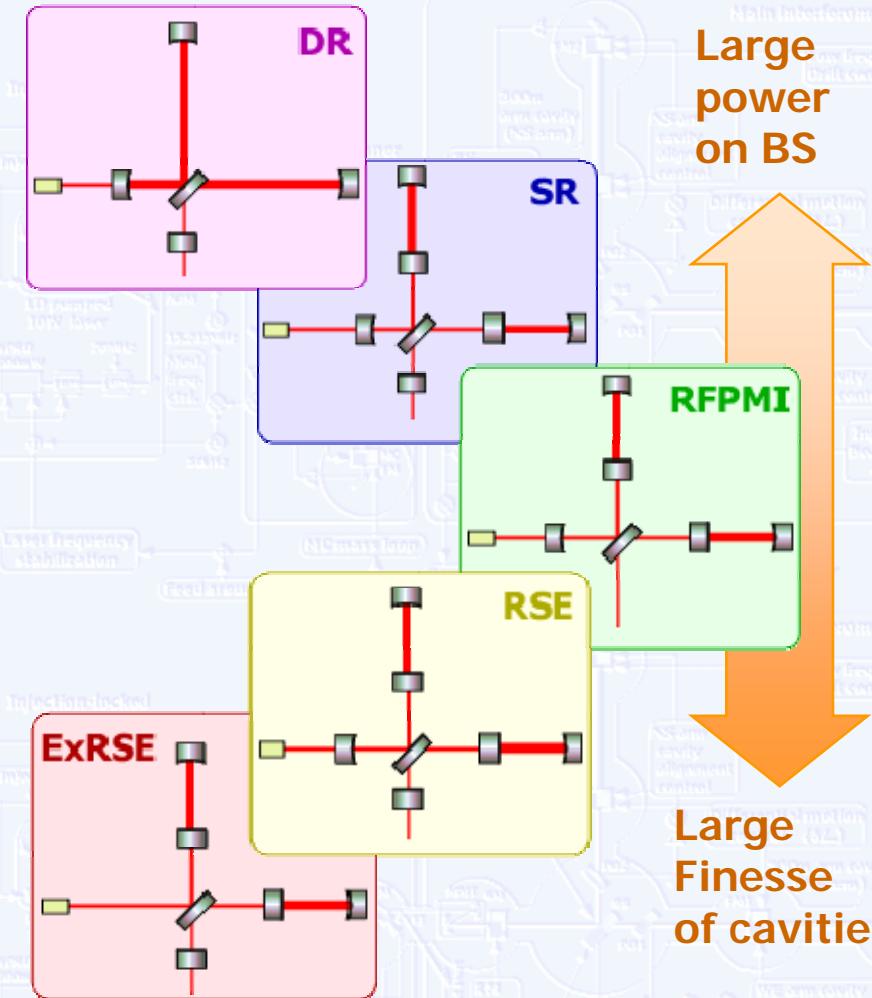


Realistic constraint

Loss in optics and interference

Simplicity of control system

Thermal problem in optics



# Main interferometer (4)

## • Merit of RSE

**High-finesse cavity and moderate PRG**

Easier to realize high power in cavities

**Smaller transmission light in optics**

**Flexible optimization for GW sources**

Independent adjustment of  
power in cavities and signal band  
Narrow-band observation (optional)

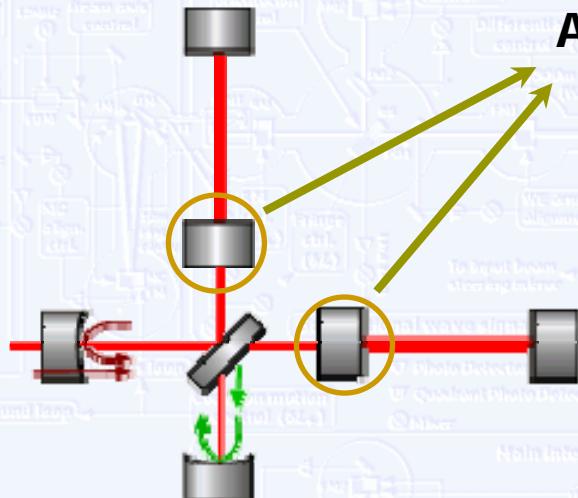
**Main reason for LCGT**

**Absorption in sapphire substrates**

Heat absorption :

$$20\text{ppm}/\text{cm} \times 15\text{ cm} = 300\text{ ppm}$$

Cooling power : 1W for each mirror



**Laser power on BS should be  
less than ~1kW (safety factor 3)**

# Main parameters

## Detector parameters

### Laser

Nd:YAG laser (1064nm)

Injection lock + MOPA

Power : **150 W**

### Main Interferometer

Broad band RSE configuration

Baseline length : **3km**

Beam Radius : **3-5cm**

Arm cavity Finesse : **1550**

Power Recycling Gain : **11**

Signal Band Gain : **15**

Stored Power : **771kW**

Signal band : **230Hz**

### Vacuum system

Beam duct diameter : **100cm**

Pressure :  **$10^{-9}$  Torr**

### Mirror

Sapphire substrate  
+ mirror coating

Diameter : **25cm**

Thickness : **15cm**

Mass : **30 kg**

Absorption Loss : **20ppm/cm**

Temperature : **20 K**

$Q = 10^8$

Loss of coating :  **$10^{-4}$**

### Final Suspension

Suspension + heat link  
with 4 Sapphire fibers

Suspension length : **40cm**

Fiber diameter : **1.5mm**

Temperature : **16K**

$Q$  of final suspension :  **$10^8$**